

THE MAGELLANIC CLOUDS HISTORY TOLD BY THEIR STAR CLUSTERS

Affiliation:

UNIVERSIDAD DE TARAPACÁ Universidad del Estado

 \star

TELESCOPE

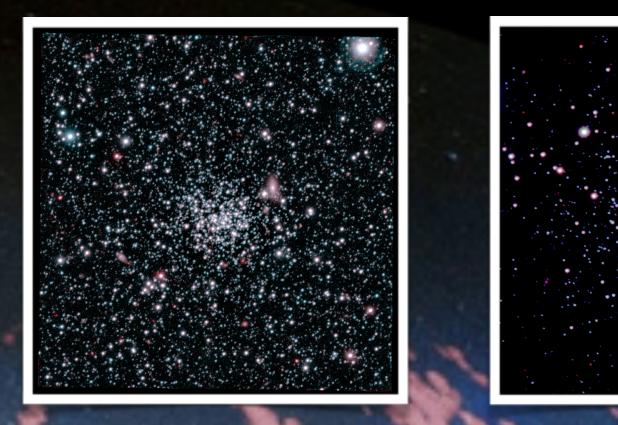
+

VISCACHA survey observations: SOAR 

Support:







Bruno Dias, IAI/UTA, CHILE

Jul 27, 2022 Gemini Science Meeting

Fondecyt Fondo Nacional de Desarrollo Científico y Tecnológico



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Affilia

VISC, survey observations: SOAR TELESCOPE







2nd GSM 2007, FOZ DO IGUAÇU, BRAZIL



Solution</t



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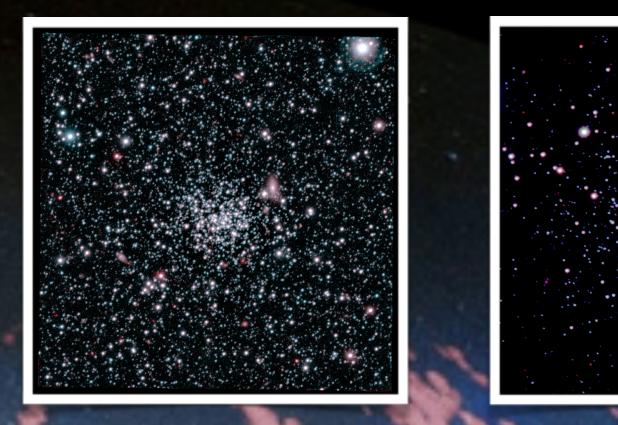
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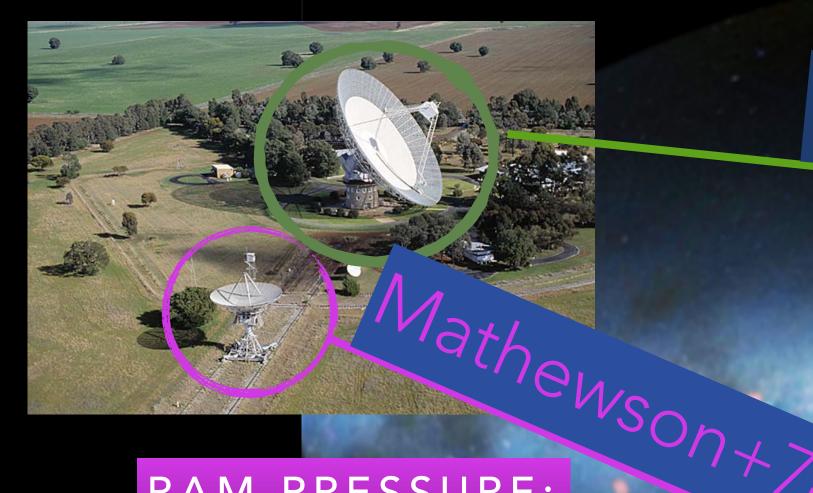


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200° MAGELLANIC STREAM (TRAILING) + LEADING ARM: RAM PRESSURE VS. TIDAL ORIGIN

Parkes, 18m, 64m





RAM PRESSURE:

ONLY GAS LEFT BEHIND

Gunn & Gott (1972) Moore & Davies (1994)

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TIDAL STRIPPING:

GAS + STARSPRESENT LEADING AND TRAILING STRUCTURES

> Broeils & van Woerden (1994) Gardiner & Noguchi (1996)









BOUND Diaz & Bekki (2012)

- LMC and SMC have been a pair, LMC and SMC have been bound to but far away from the Milky Way the Milky Way
- Only recently (~2 Gyr ago) LMC+SMC became a pair

Orbital period 1.5-2 Gyr

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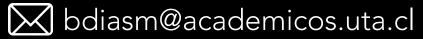
- Only recently (~2 Gyr ago) LMC+SMC arrived to the MW vicinity
- Hyperbolic orbit (1st infall) or 2nd passage (orbital period ~6 Gyr)







DO THE MODELS REPRODUCE THE OBSERVATIONAL EVIDENCE?

















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DUCE THE NCE?	









Position*, morphology, kinematic <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et</u>

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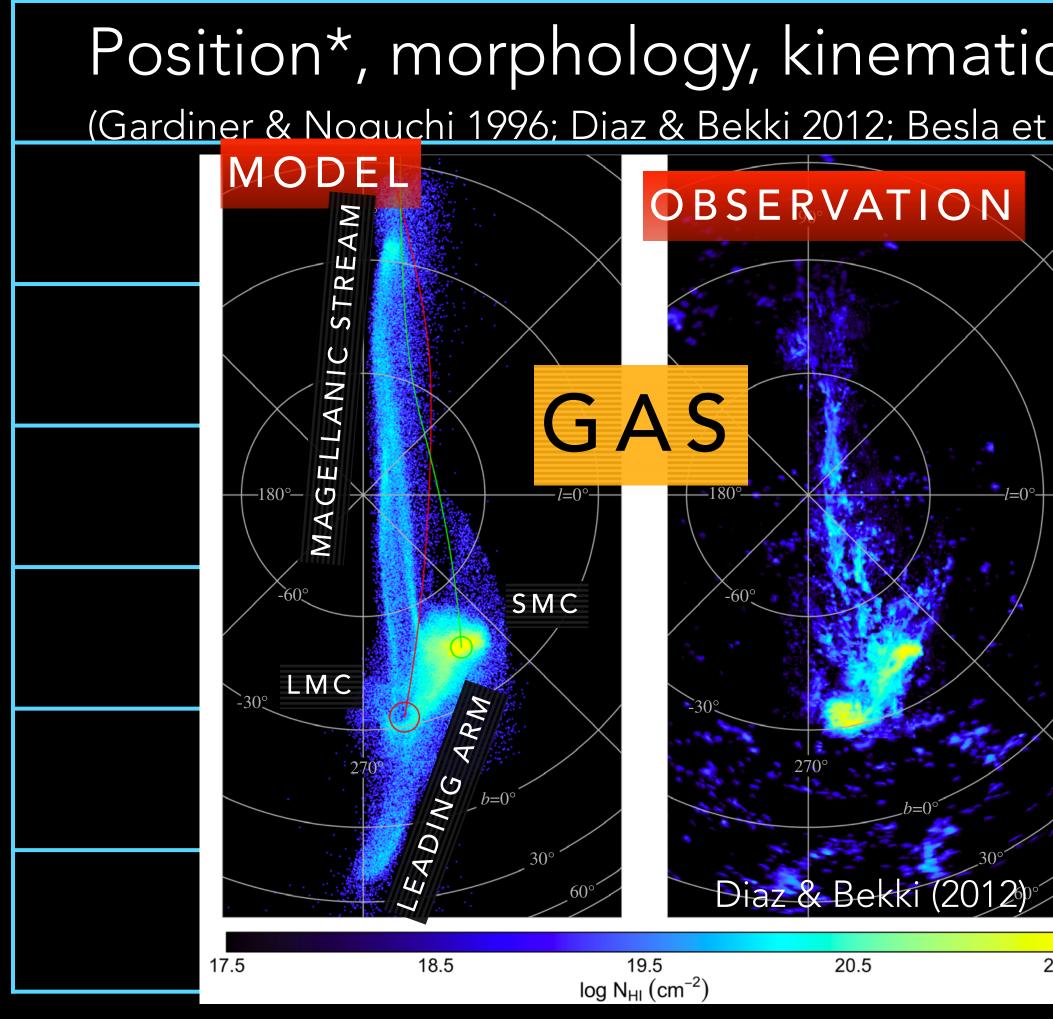
DUCE THE NCE?	
cs Mag. Stream t al. 2010, 2012)	











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DUCE THE NCE?	
cs Mag. Stream t al. 2010, 2012)	
21.5	

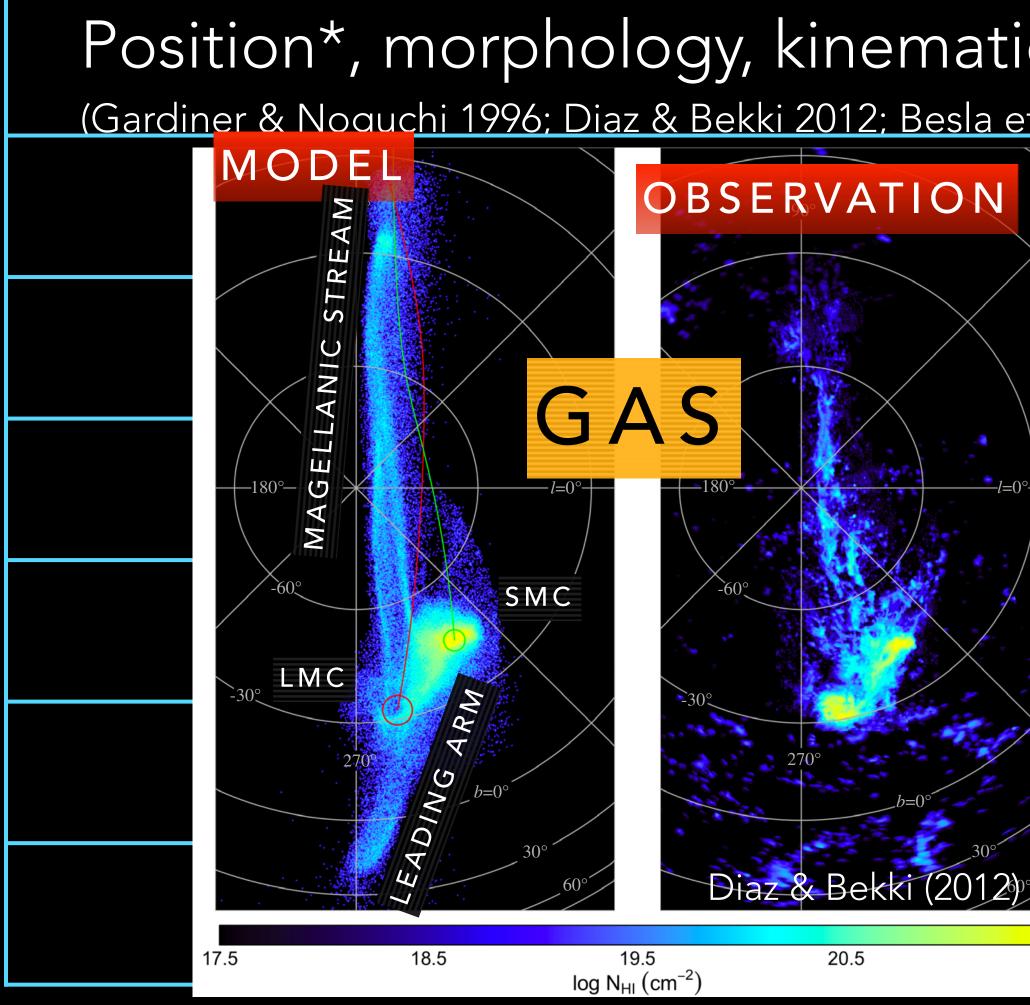








DO THE MODELS REPRODUCE THE OBSERVATIONAL EVIDENCE? Position*, morphology, kinematics Mag. Stream <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et al. 2010, 2012)</u> MODEL OBSERVATION OBSERVATION STRE GELLANIC GAS Nidever et al. (2010) Model 2 GAS MODEL SMC LMC ARM DING Stream Ē Diaz & Bekki (2012)° Besla et al. (2012) 20.5 17.5 18.5 19.5 21.5 20 17 18 19 21 $\log N_{HI} \, (cm^{-2})$ Log (N_{gos} cm⁻²) AstroBDias AstroBDias www.astro.iag.usp.br/~viscacha o astrobdias



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Position*, morphology, kinematic <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et</u>

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DUCE THE NCE?	
cs Mag. Stream t al. 2010, 2012)	









Position*, morphology, kinematic <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et</u>

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DUCE THE NCE?		
cs Mag. Stream _{t al.} 2010, 2012)	YES	







Position*, morphology, kinematic <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et</u>

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DUCE THE NCE?		
cs Mag. Stream _{t al.} 2010, 2012)	YES	YES









Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>		

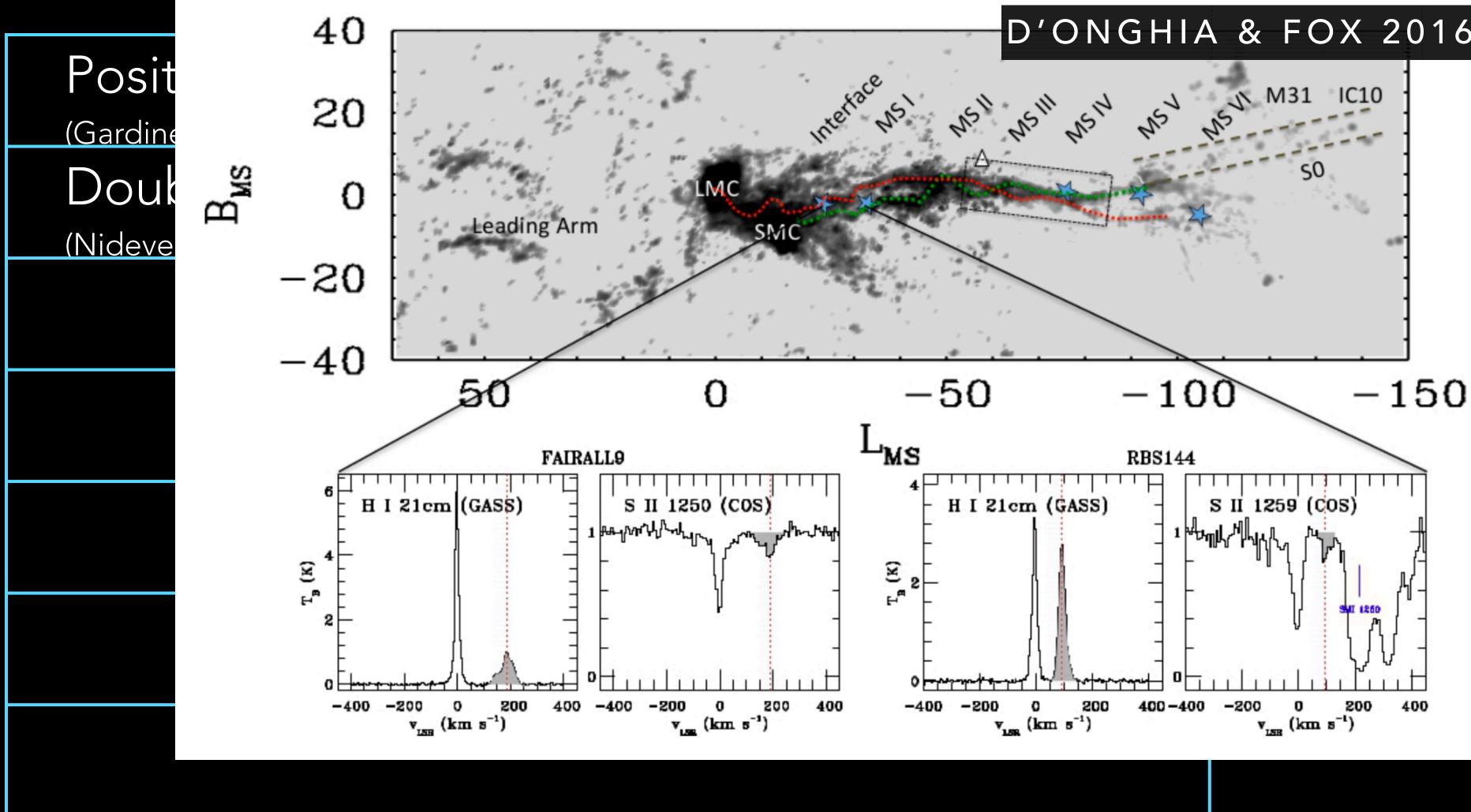








DO THE MODELS REPRODUCE THE OBSERVATIONAL EVIDENCE?



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D'ONGHIA & FOX 2016









Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>		









Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	









Position*, morphology, kinematic (Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream <u>(Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201</u>

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?









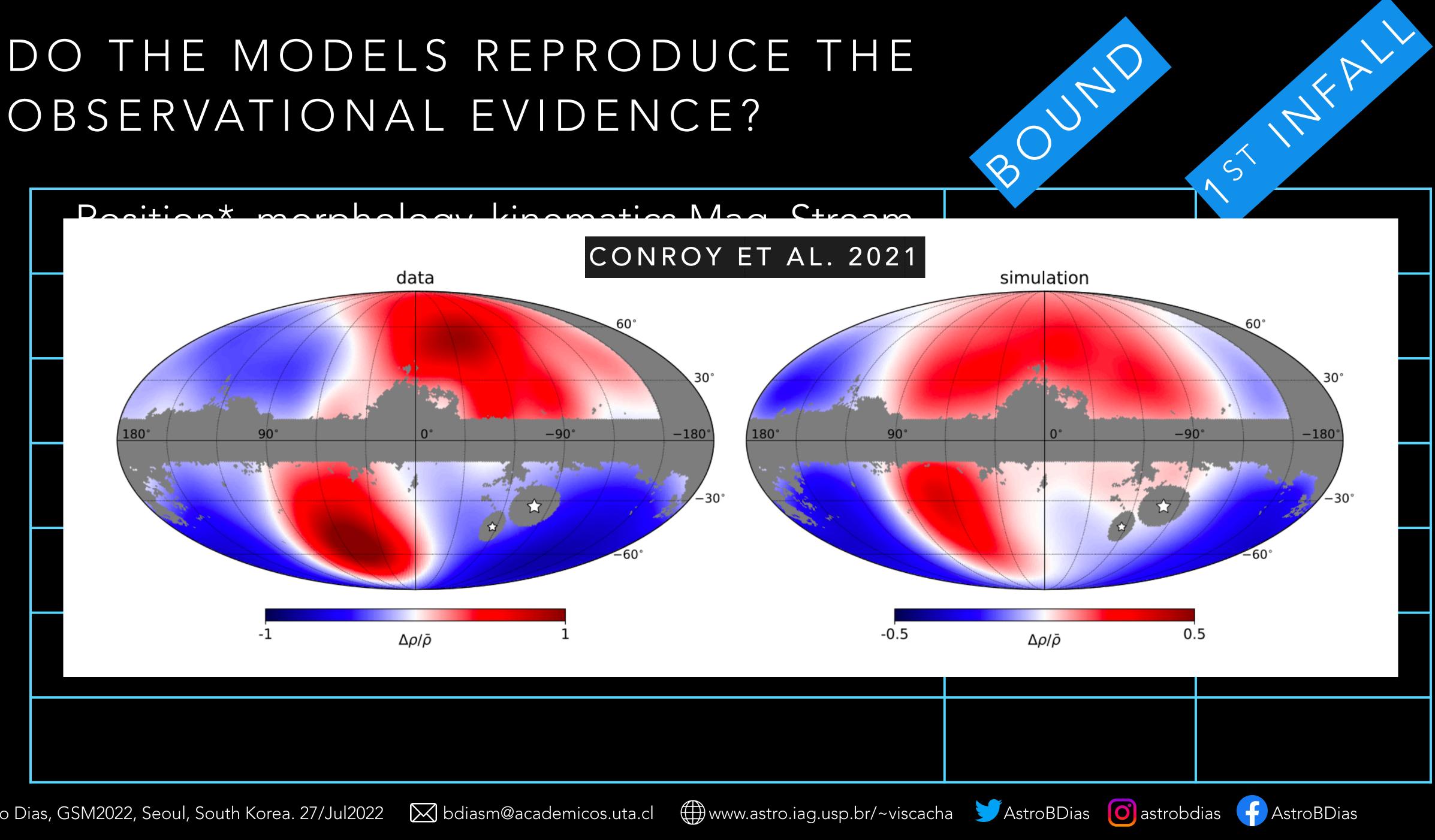
DUCE THE NCE?		
cs Mag. Stream	YES	YES
<u>12: D'Onghia & Fox 2016)</u> verdensity	YES	NO?











DUCE THE NCE?		
cs Mag. Stream	YES	YES
<u>12: D'Onghia & Fox 2016)</u> verdensity	YES	NO?









DUCE THE NCE?		
cs Mag. Stream	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	ΝΟ	









DUCE THE NCE?		
cs Mag. Stream	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	ΝΟ	YES









Position*, morphology, kinematic <u>(Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et</u> Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021)

Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u>

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	ΝΟ	YES
S = LMC? 21; Hammer et al. 2021)		

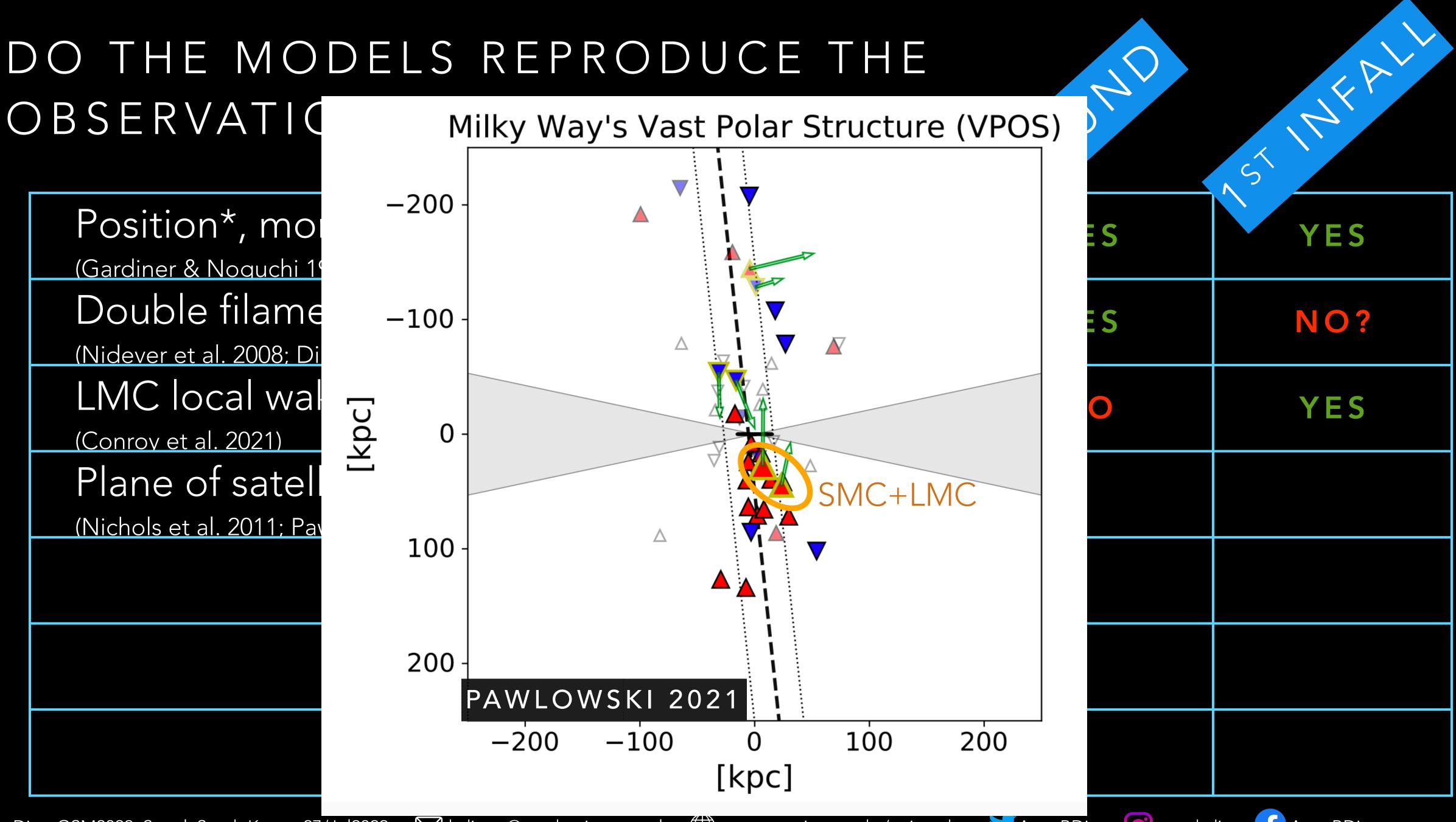












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Position*, morphology, kinematic <u>(Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et</u> Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021)

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DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	ΝΟ	YES
S = LMC? 21; Hammer et al. 2021)		











Position*, morphology, kinematic <u>(Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et</u> Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021)

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DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	

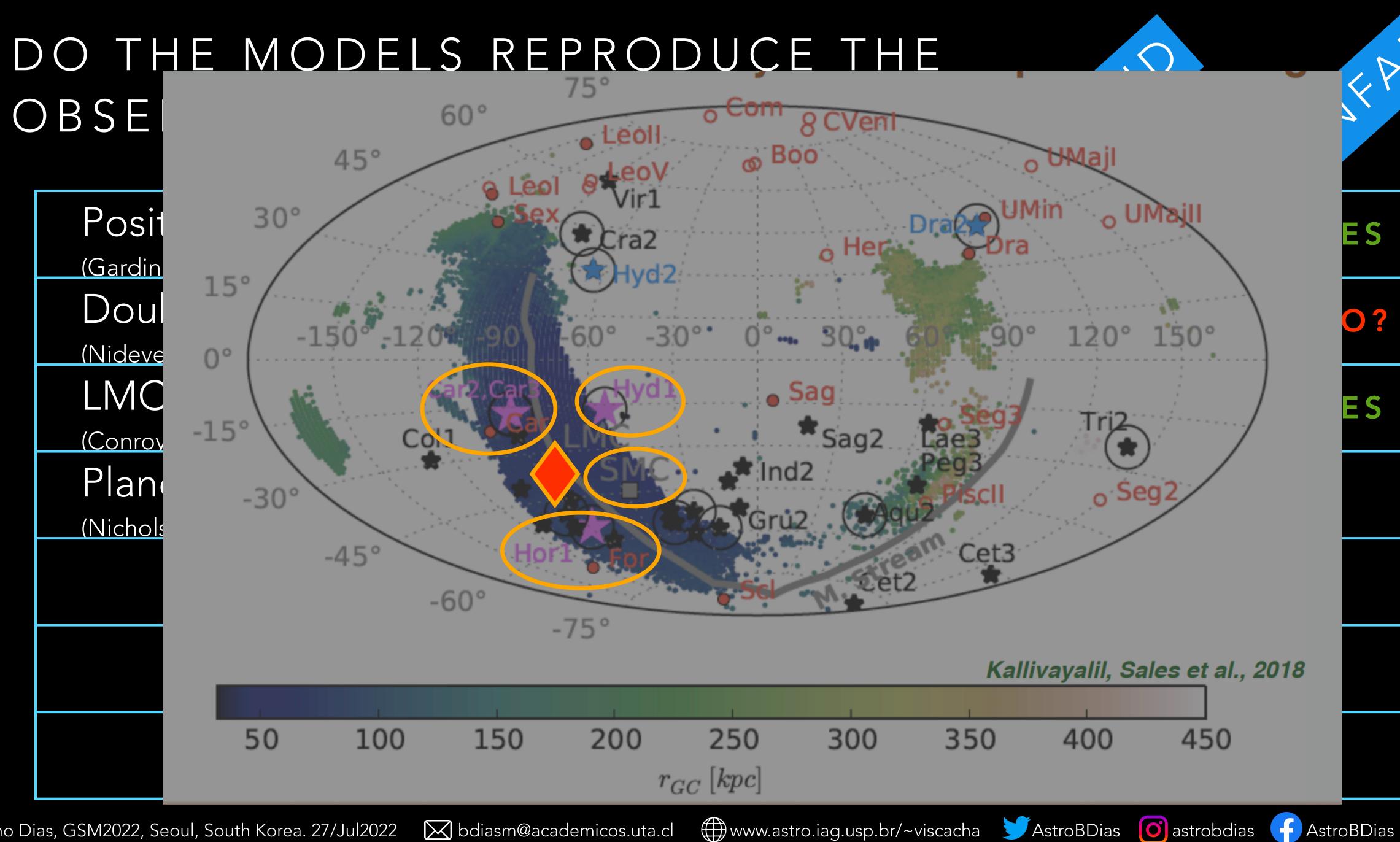




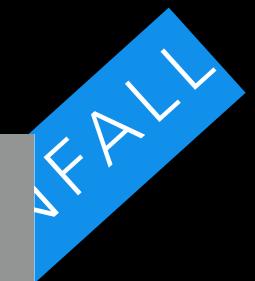








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Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u>

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	











Position*, morphology, kinematic <u>(Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et</u> Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021)

Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u>

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	ΝΟ	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??











Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	N O ?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
<u>al. 2019, 2020)</u>		









Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	









Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 2021</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O









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Position*, morphology, kinematic <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et</u> Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS (Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202 What is the LMC mass? (Kallivavalil et al. 2013; Guglielmo et al. 2014; Erkal et a How is the MW potential? (Besla 2007; Nidever et al. 2008; Diaz & Bekki 2012)

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O









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Position*, morphology, kinematic <u>(Gardiner & Noquchi 1996; Diaz & Bekki 2012; Besla et</u> Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS (Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 2021 What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a How is the MW potential? (Besla 2007; Nidever et al. 2008; Diaz & Bekki 2012)

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	ΝΟ	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O
	SIMPLE (ISOTHERMAL SPHERE)	









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Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a How is the MW potential? <u>(Besla 2007; Nidever et al. 2008; Diaz & Bekki 2012)</u>

DUCE THE NCE?		
cs Mag. Stream t al. 2010, 2012)	YES	YES
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O
	SIMPLE (ISOTHERMAL SPHERE)	REALIST (NFW)











DO THE MODELS REPRO OBSERVATIONAL EVIDEN

Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a How is the MW potential? <u>(Besla 2007; Nidever et al. 2008; Diaz & Bekki 2012)</u> What is the source and precision (Kallivayalil et al. 2006, Vieira et al. 2010, Niederhofer e

DUCE THE NCE?			
cs Mag. Stream t al. 2010, 2012)	YES	YES	
<u>12; D'Onghia & Fox 2016)</u>	YES	NO?	
verdensity	NO	YES	
S = LMC? 21; Hammer et al. 2021)	YES??	YES??	
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O	
	SIMPLE (ISOTHERMAL SPHERE)	REALIST (NFW)	
n of the PMs et al. 2018, Helmi et al. 2018)			
cl			





DO THE MODELS REPRO OBSERVATIONAL EVIDEN

Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a How is the MW potential? <u>(Besla 2007; Nidever et al. 2008; Diaz & Bekki 2012)</u> What is the source and precision (Kallivayalil et al. 2006, Vieira et al. 2010, Niederhofer et

DUCE THE NCE?		
cs Mag. Stream _{t al.} 2010, 2012)	YES	YES
12; D'Onghia & Fox 2016)	YES	NO?
verdensity	NO	YES
S = LMC? 21; Hammer et al. 2021)	YES??	YES??
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O
	SIMPLE (ISOTHERMAL SPHERE)	REALIST (NFW)
n of the PMs It al. 2018, Helmi et al. 2018)	GROUND <0.3 mas/yr	
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DO THE MODELS REPRO OBSERVATIONAL EVIDEN

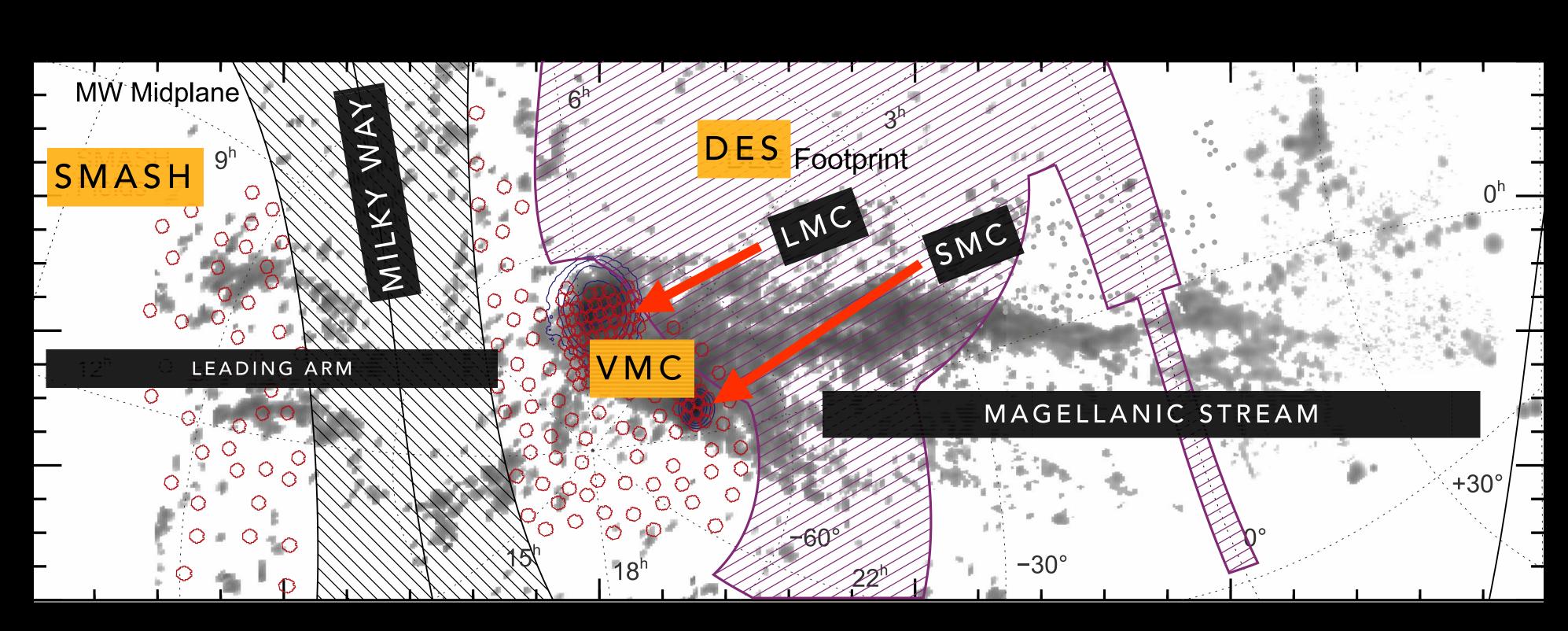
Position*, morphology, kinematic (Gardiner & Noguchi 1996; Diaz & Bekki 2012; Besla et Double filament in Mag. Stream (Nidever et al. 2008; Diaz & Bekki 2011; Besla et al. 201 LMC local wake and Northern ov (Conroy et al. 2021) Plane of satellites: origin of VPOS <u>(Nichols et al. 2011; Pawlowski 2018, 2021; Li et al. 202</u> What is the LMC mass? (Kallivayalil et al. 2013; Guglielmo et al. 2014; Erkal et a How is the MW potential? <u>(Besla 2007; Nidever et al. 2008; Diaz & Bekki 2012)</u> What is the source and precision (Kallivayalil et al. 2006, Vieira et al. 2010, Niederhofer e[.]

DUCE THE NCE?			
cs Mag. Stream t al. 2010, 2012)	YES	YES	
12; D'Onghia & Fox 2016)	YES	NO?	
verdensity	NO	YES	
S = LMC? 21; Hammer et al. 2021)	YES??	YES??	
al. 2019, 2020)	LIGHT (~10 ¹⁰ M _O)	HEAVY (~10 ¹¹ M _O	
	SIMPLE (ISOTHERMAL SPHERE)	REALIST (NFW)	
n of the PMs et al. 2018, Helmi et al. 2018)	GROUND <0.3 mas/yr		
cl			





MANY SURVEYS HAVE BEEN LOOKING FOR THE STELLAR COUNTERPART OF THE TIDAL EFFECTS CAUSED BY THE ORBITS OF THE SMC/LMC



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Background image: Nidever et al. (2017)

VISCACHA

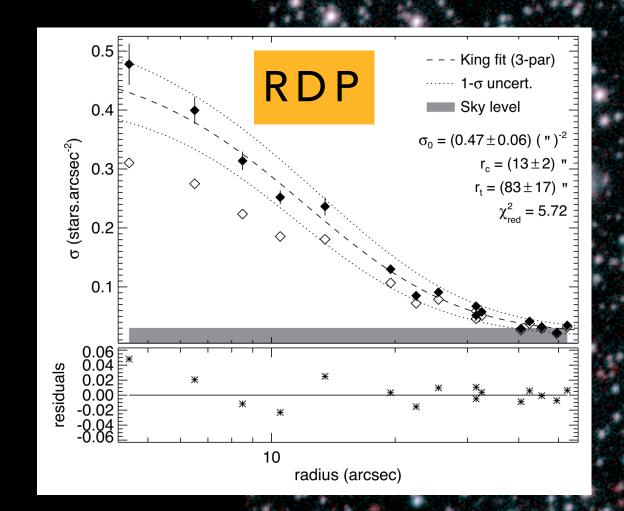


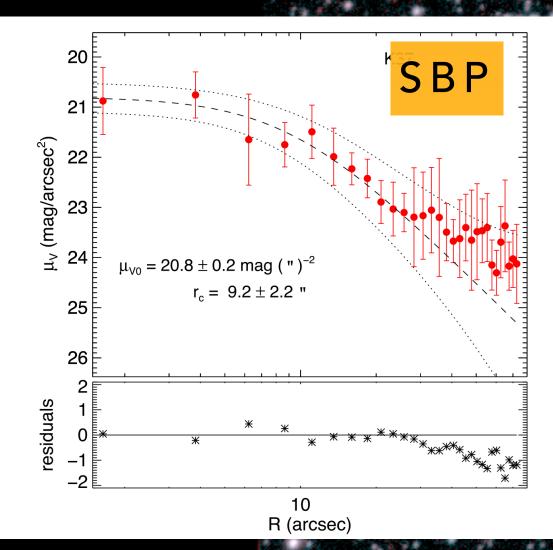
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WHY STAR



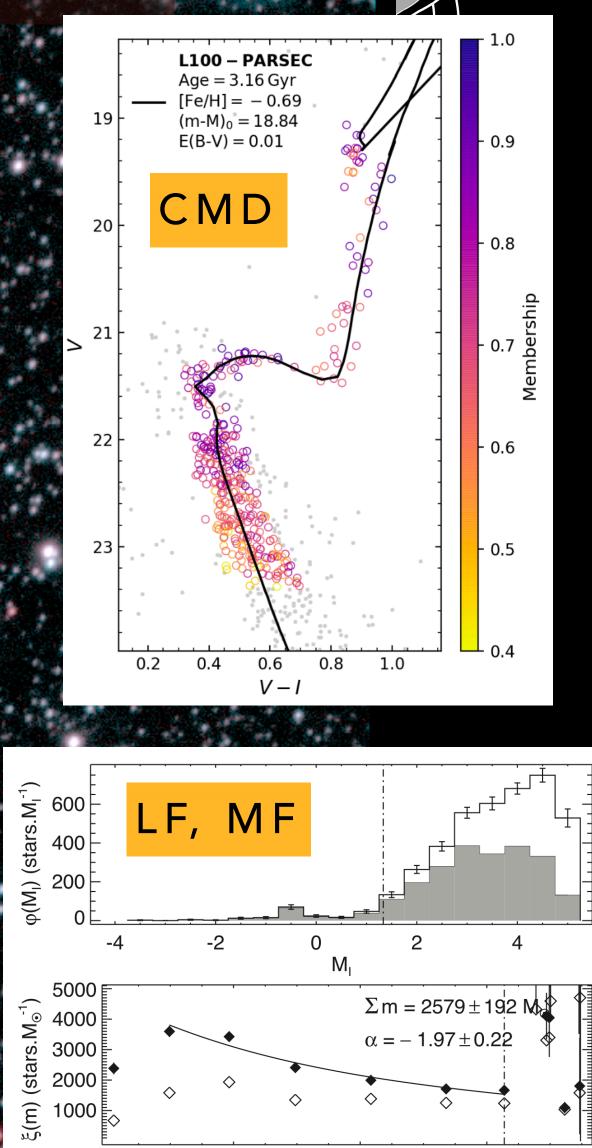


1.0

Bruno Dias, GSM2022, Seoul, South Ko

STAR CLUSTERS?

30″



astropolas

1.4

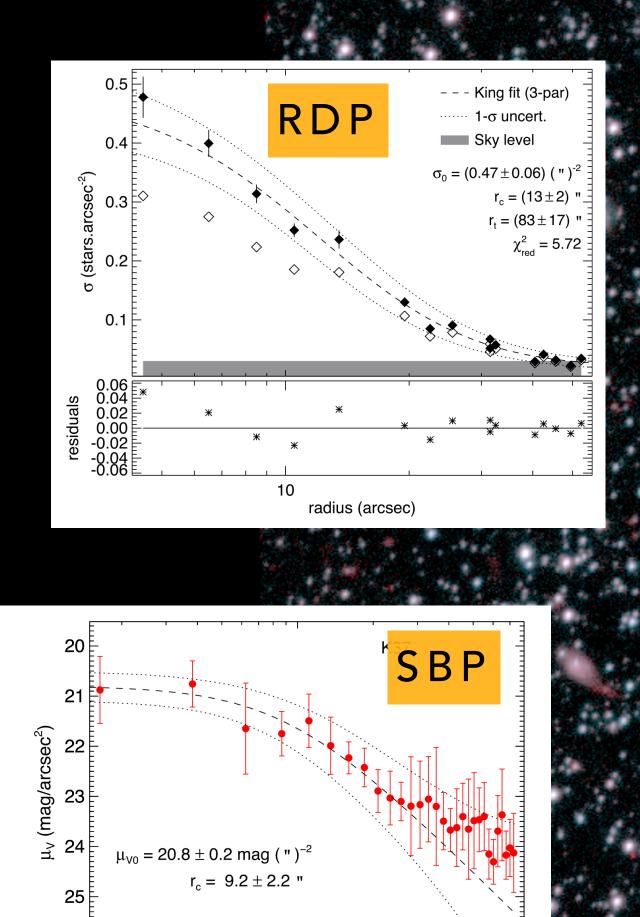
1.2 m (M_☉)

1.0

Maia, Dias et al. (2019)

AstroBDias

WHY STAR CLUSTERS?



10 R (arcsec)

1.000

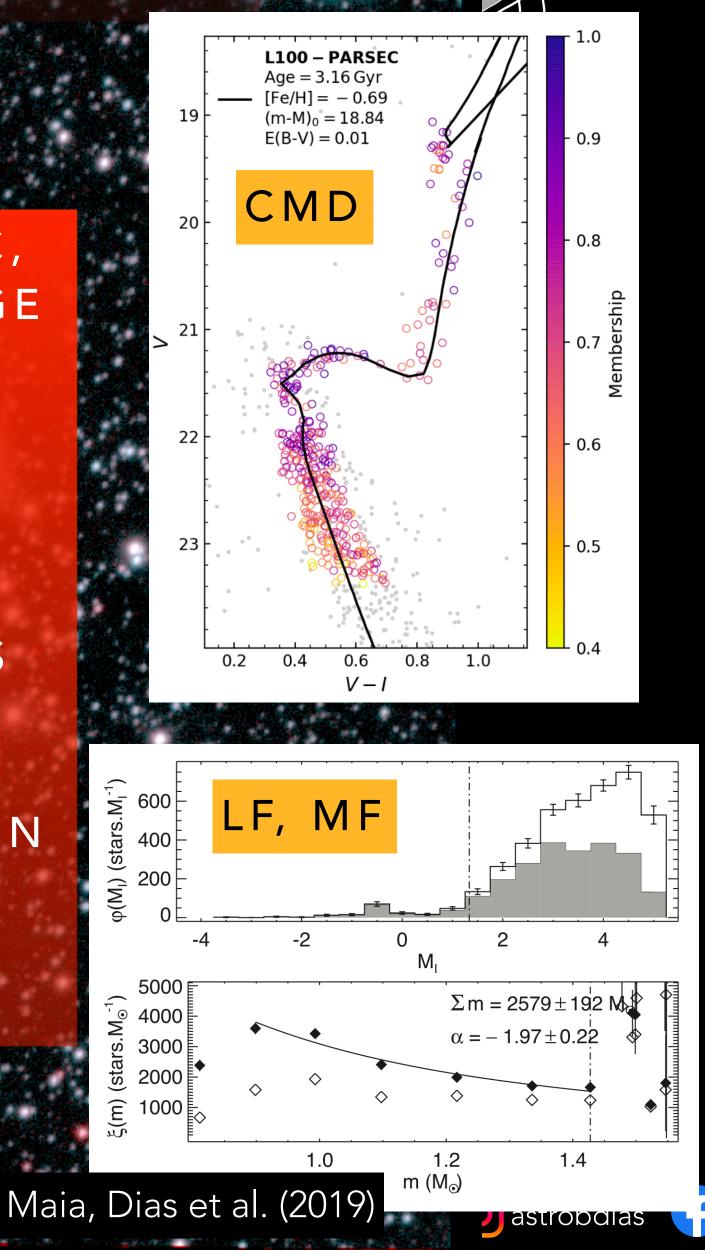
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26 İ

residuals 1 0 1-

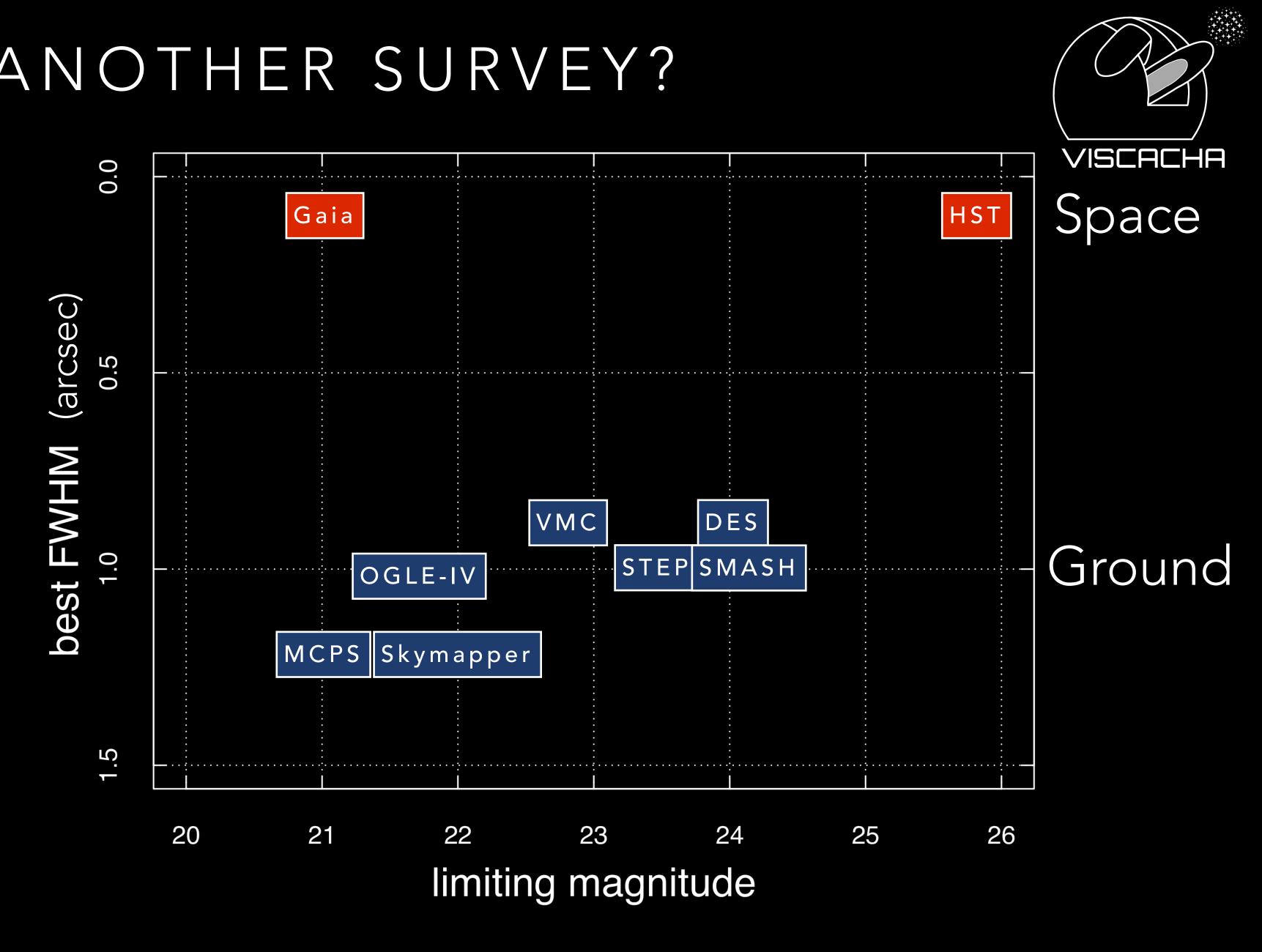
3D MAP OF SMC, LMC, AND BRIDGE WITH: AGE METALLICITY REDDENING CORE RADIUS TIDAL RADIUS ELLIPTICITY TOTAL MASS MASS FUNCTION DISSOLUTION eMSTO

30"



AstroBDias

WHY ANOTHER SURVEY?



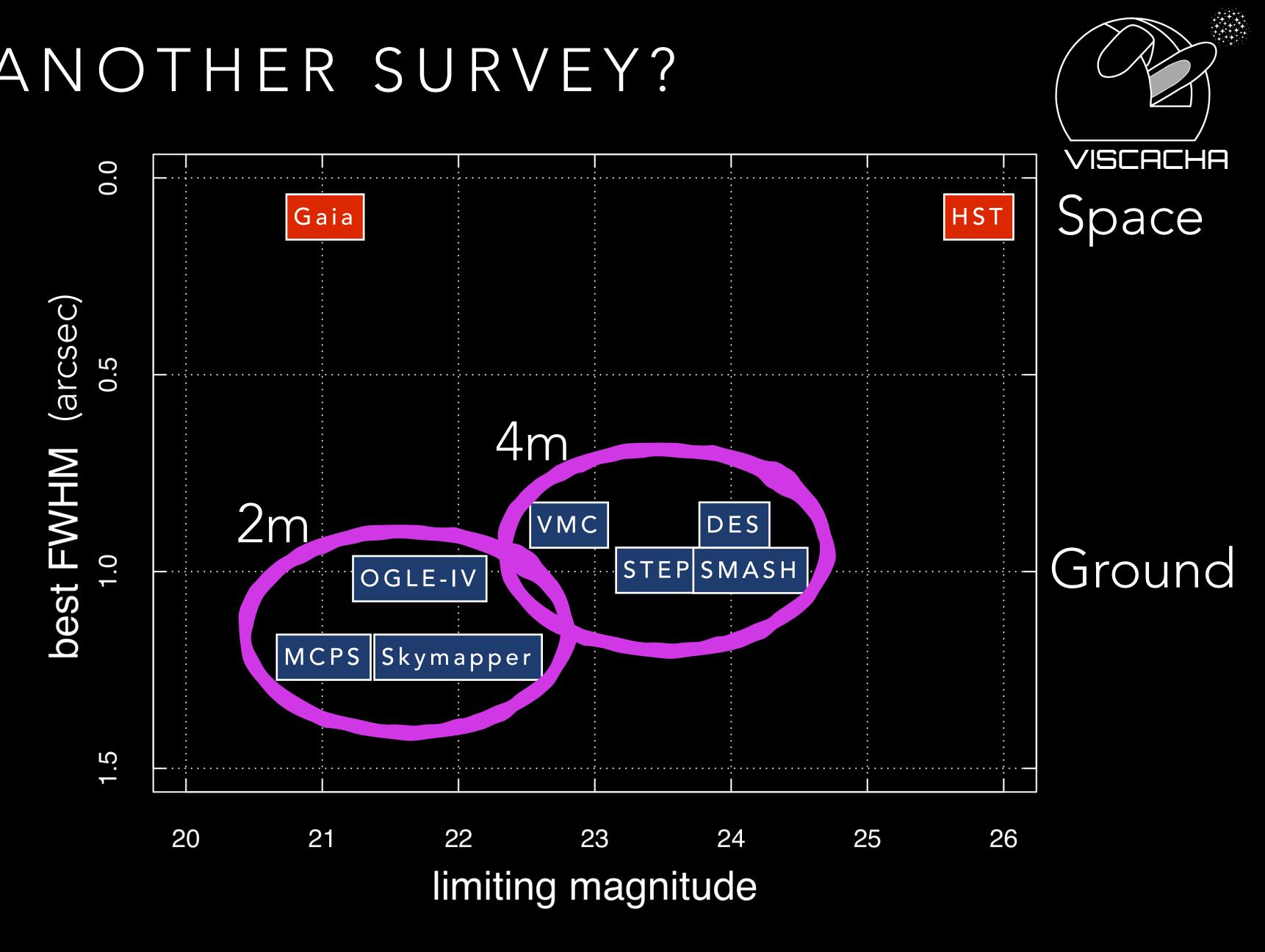
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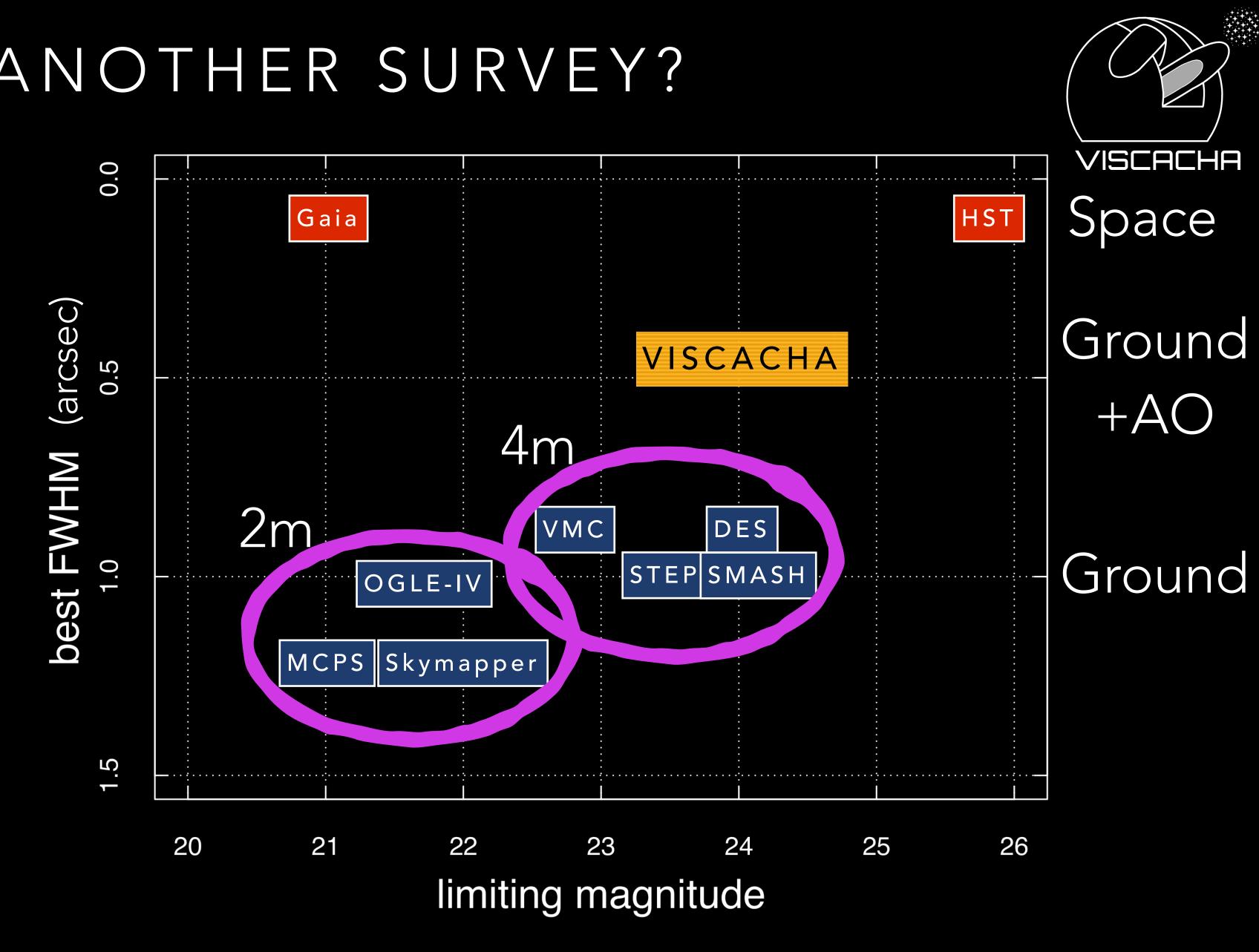
WHY ANOTHER SURVEY?



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WHY ANOTHER SURVEY?

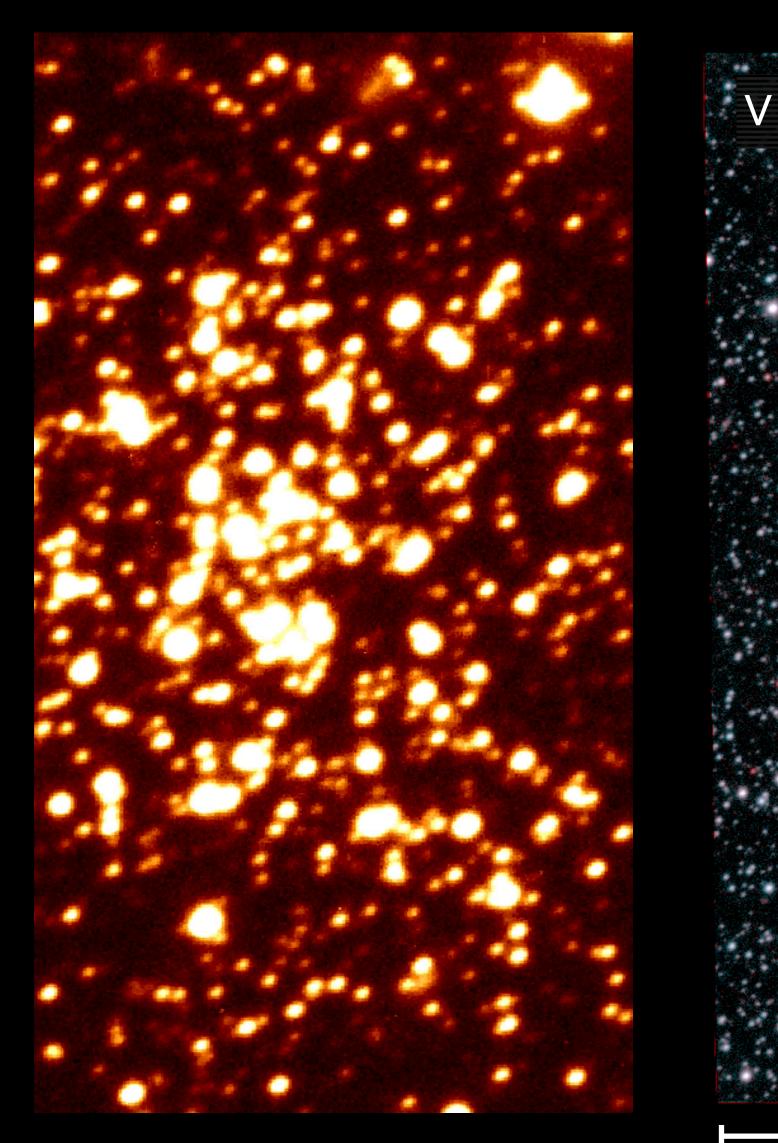


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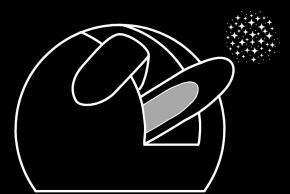


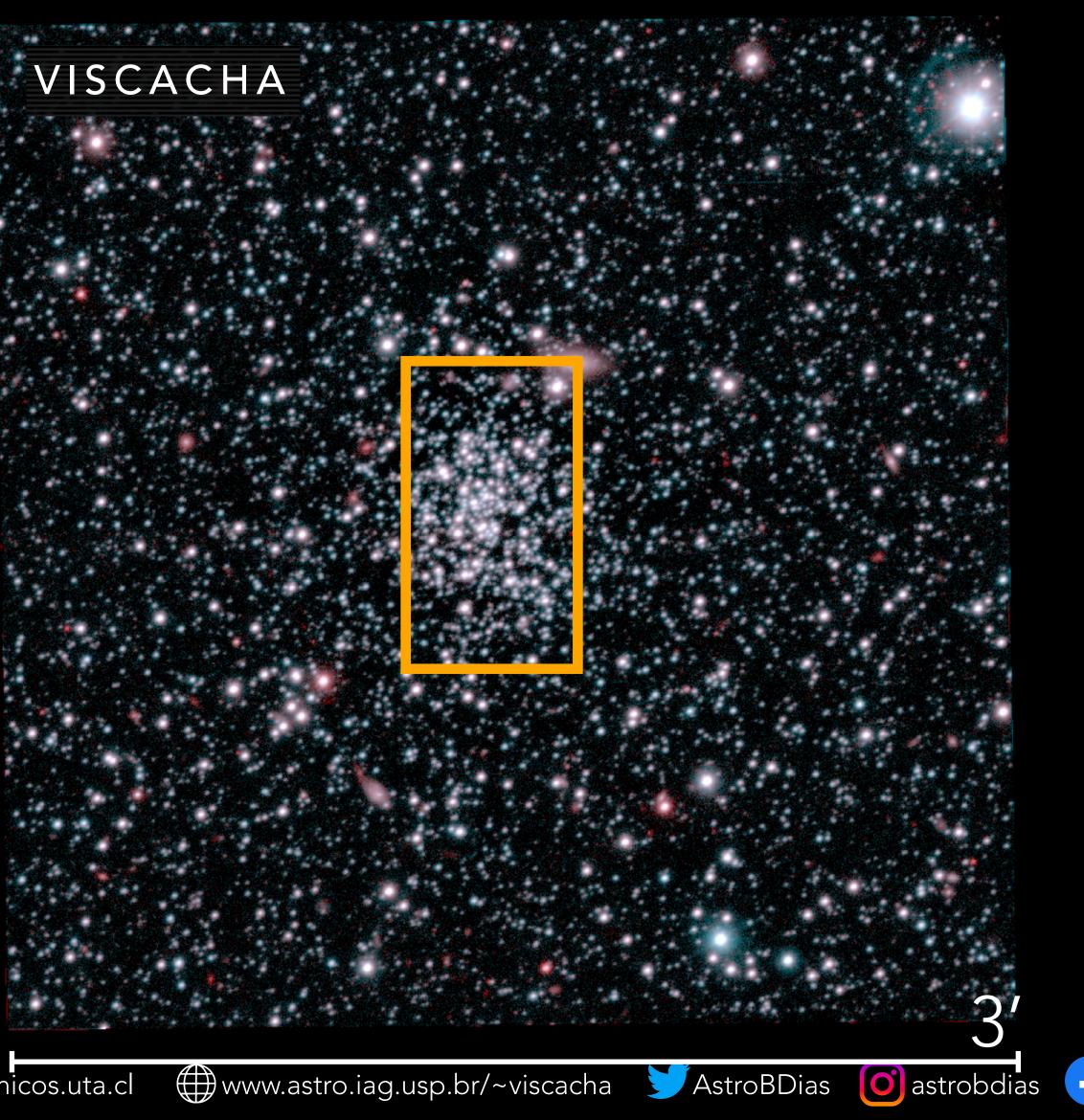


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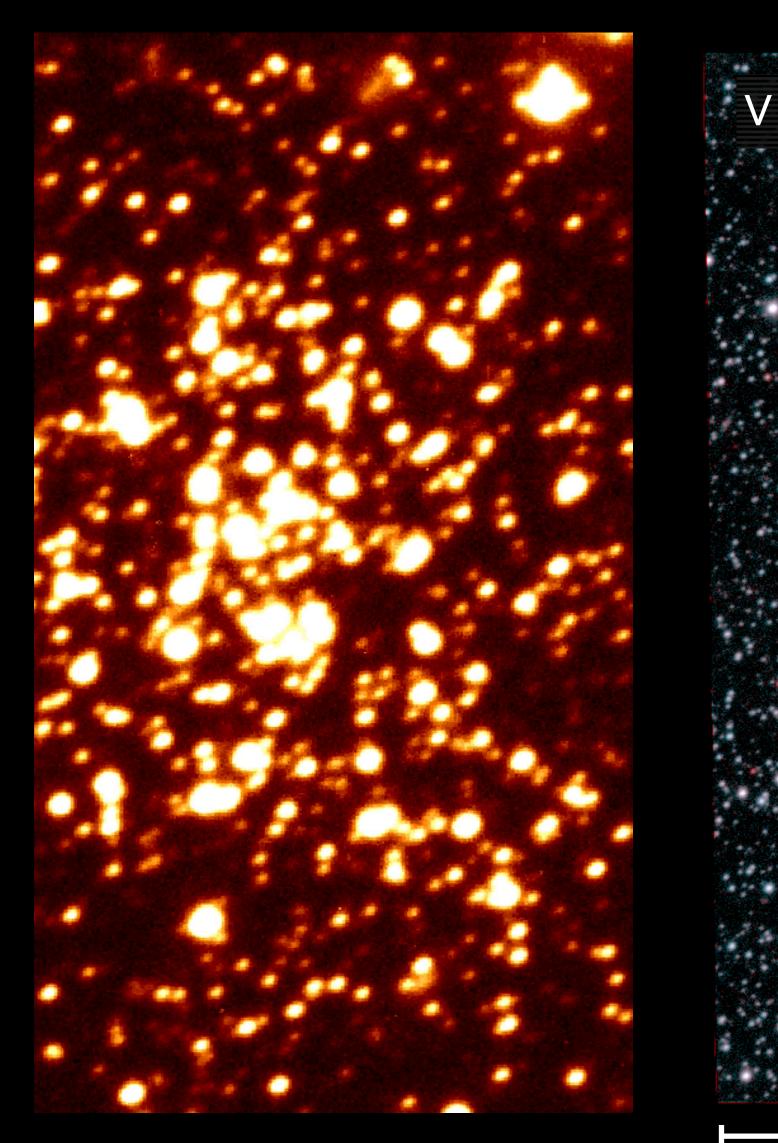
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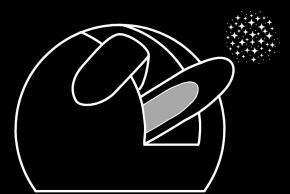


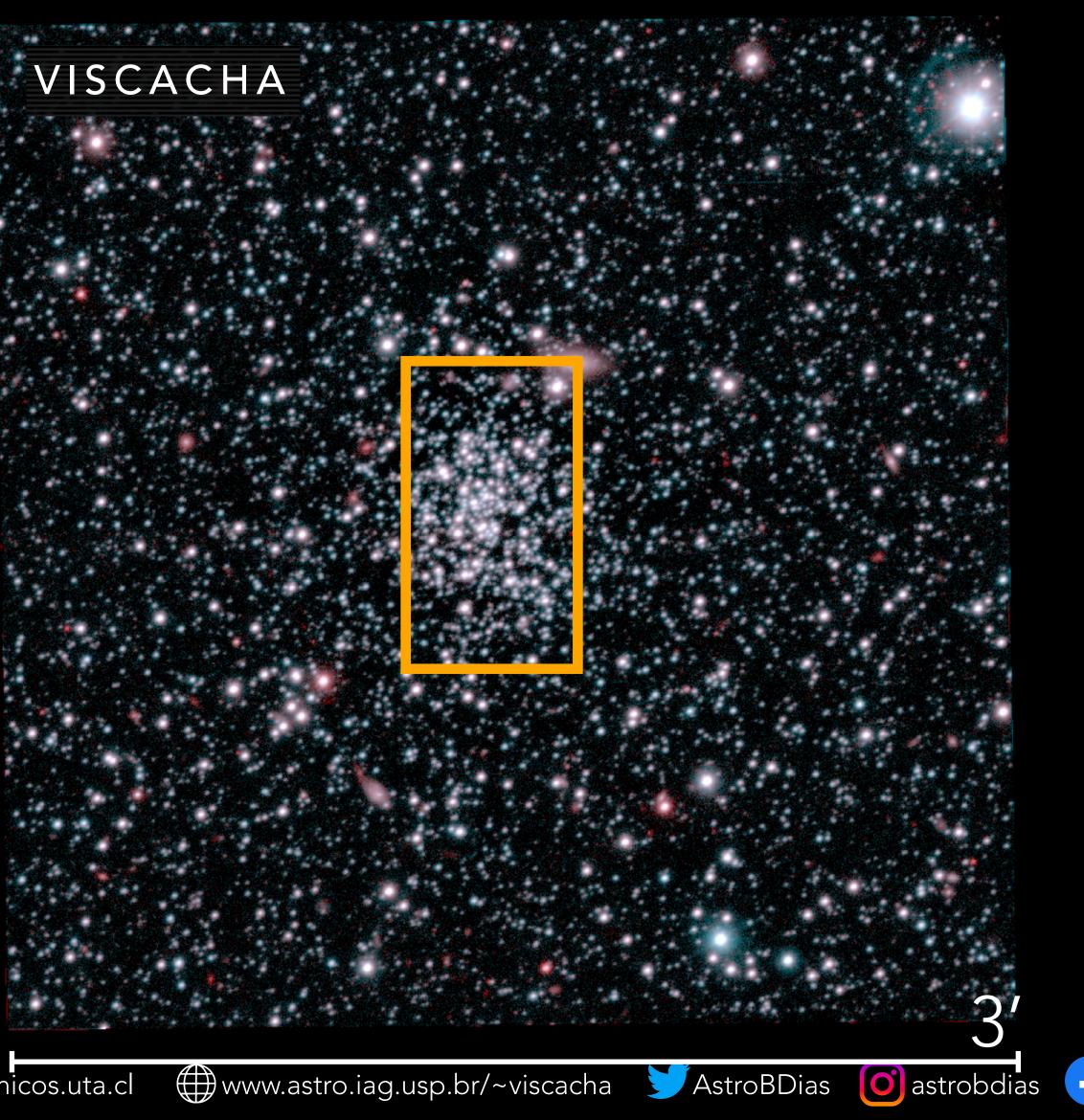


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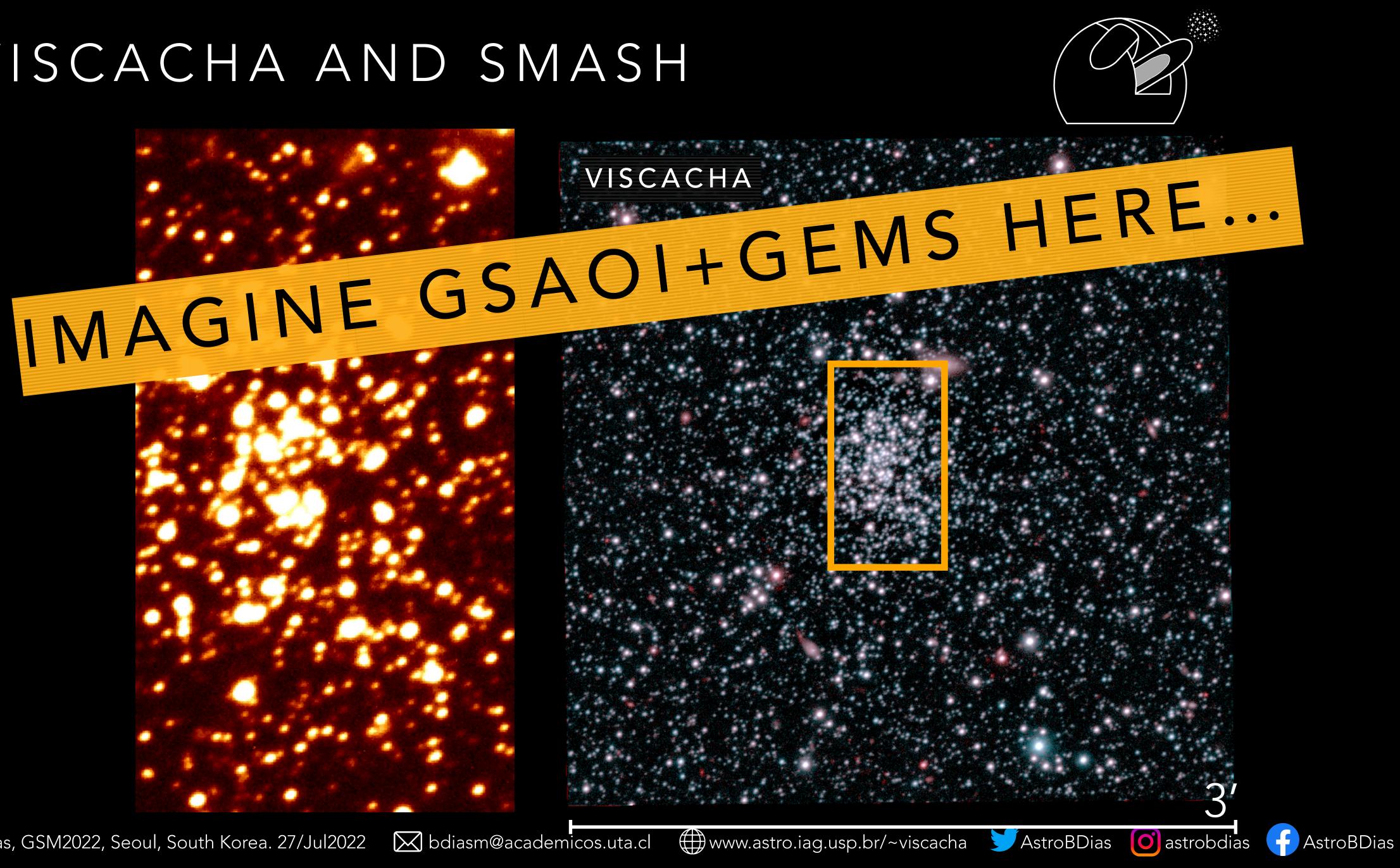


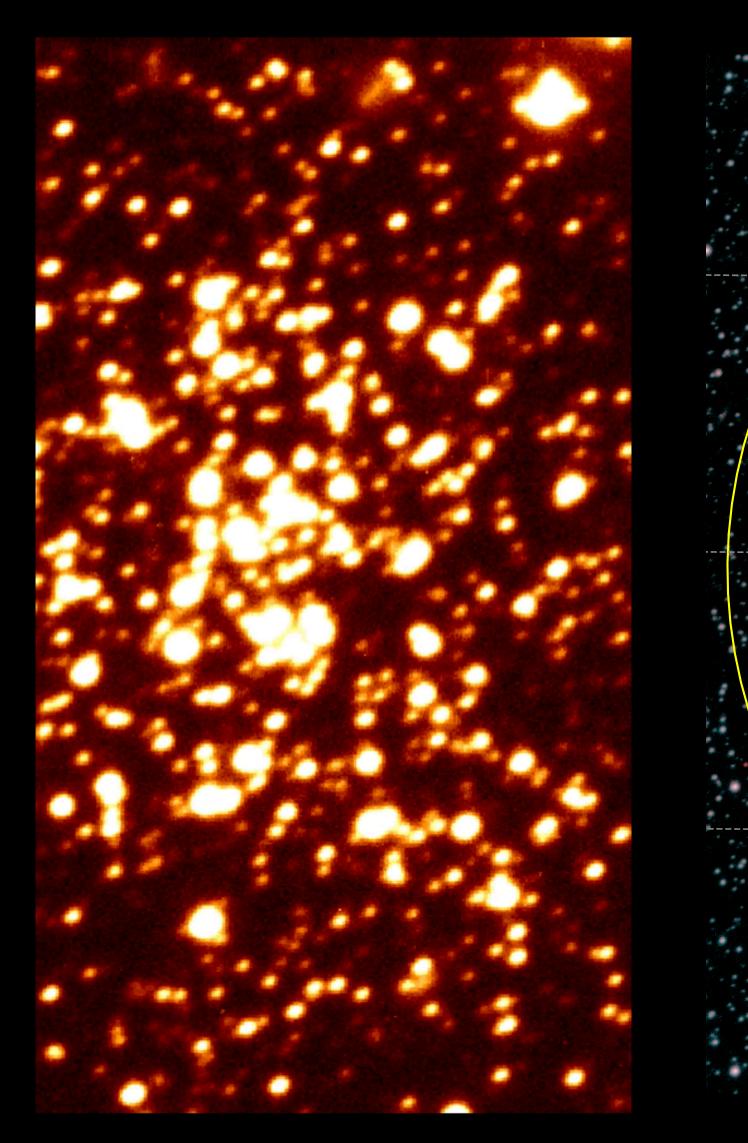






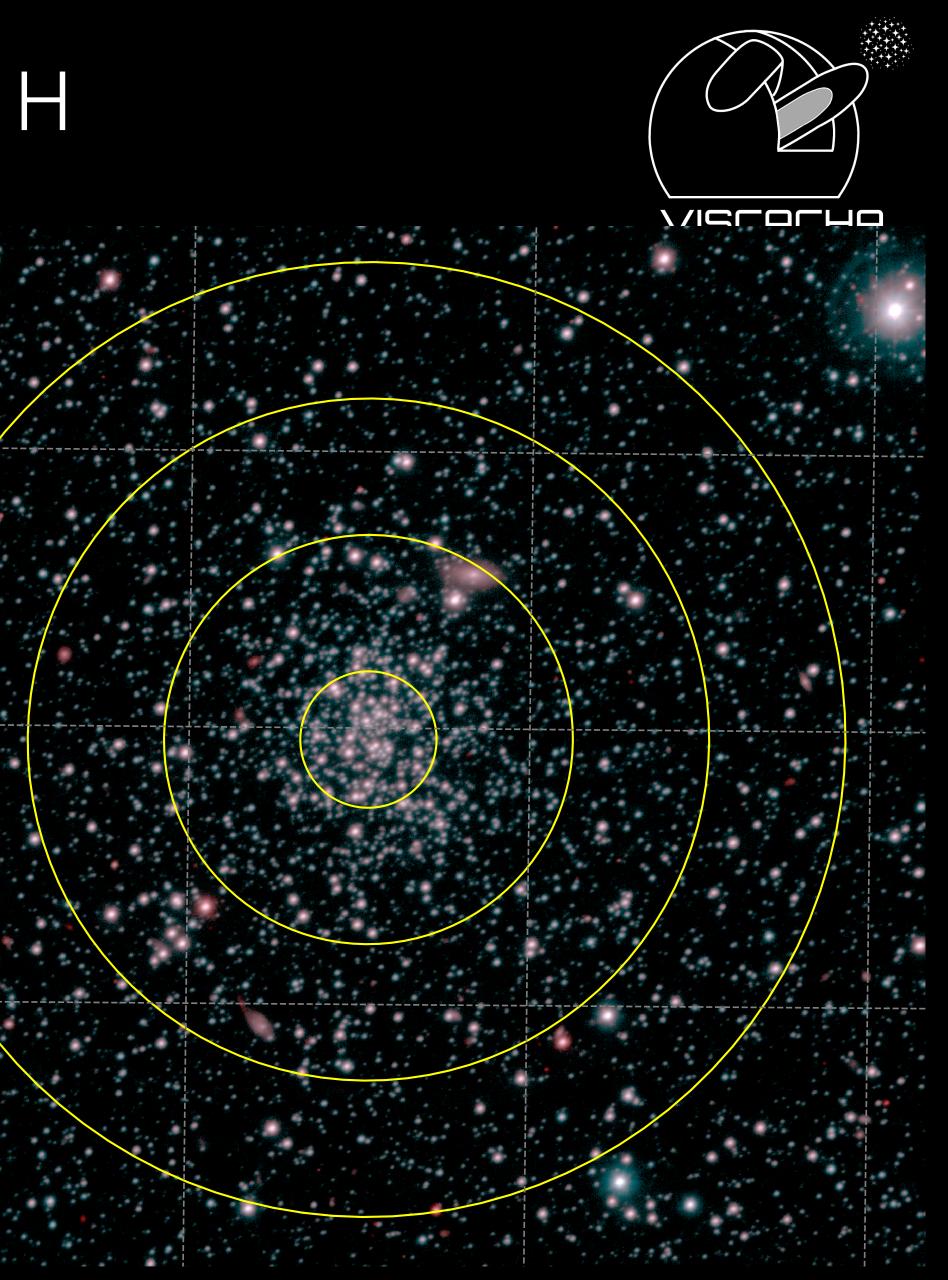
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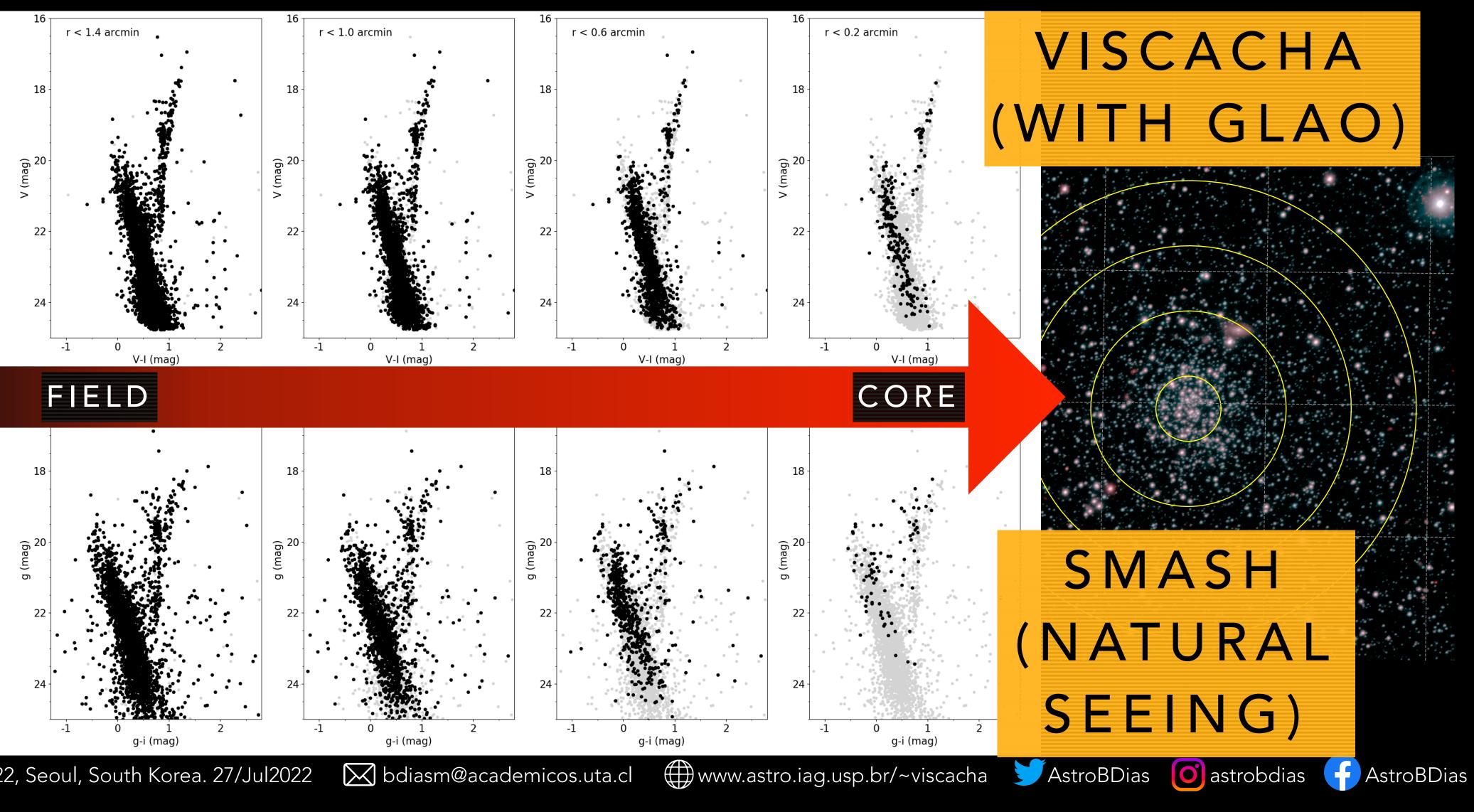
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3x10⁵ 512h (250h Sci) point sources 80% 400Gb (20Gb 2025 reduced) VISCACHA .200 images 2015 - 2021 (reduced) 215 clusters

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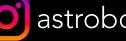
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40+ members 20+ univ.+obs. 5 countries

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3x10⁵ 512h point (250h Sci) sources 80% 400Gb (20Gb 2025 reduced) VISCACHA .200 images 2015 - 2021 (reduced) 215 clusters

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40+ members 20+ univ.+obs. 5 countries

To map the Mag. Clouds in 3D using star clusters

To discover star clusters under dissolution process

To unveil the chemical and dynamical evolution of SMC/LMC

To find LMC clusters with ages between 3-10x10⁹years

> To identify clusters with multiple stellar populations

and more!









200



512h (250h Sci)





.200 images (reduced)

215 clusters

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Credit: NOIRLab

問題とは見



40+ members 20+ univ.+obs. 5 countries







El?



VISCACHA SPECTROSCOPIC FOLLOW-UP + GAIA PROPER MOTIONS



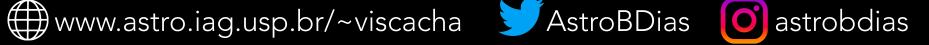
Chile (PI: Dias et al.) + Brazil (Kerber et al.) + Argentina (Parisi et al.)

Joint project: 2018,2019,2020,2021

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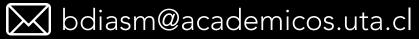
AstroBDias



VISCACHA SPECTROSCOPIC FOLLOW-UP/ SEE POSTER DE BORTOLI ET AL. + GAIA PROPER MOTIONS SEE POSTER DIAS & PARISI Call triplet spectroscopy: GMOS Metallicities Radial velocities tlux Normalized 0.6 0.8 0.4 8500 8600 8550 8650 λ (A)



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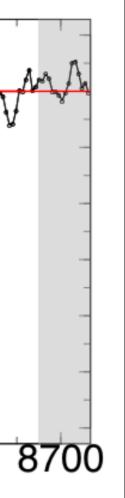




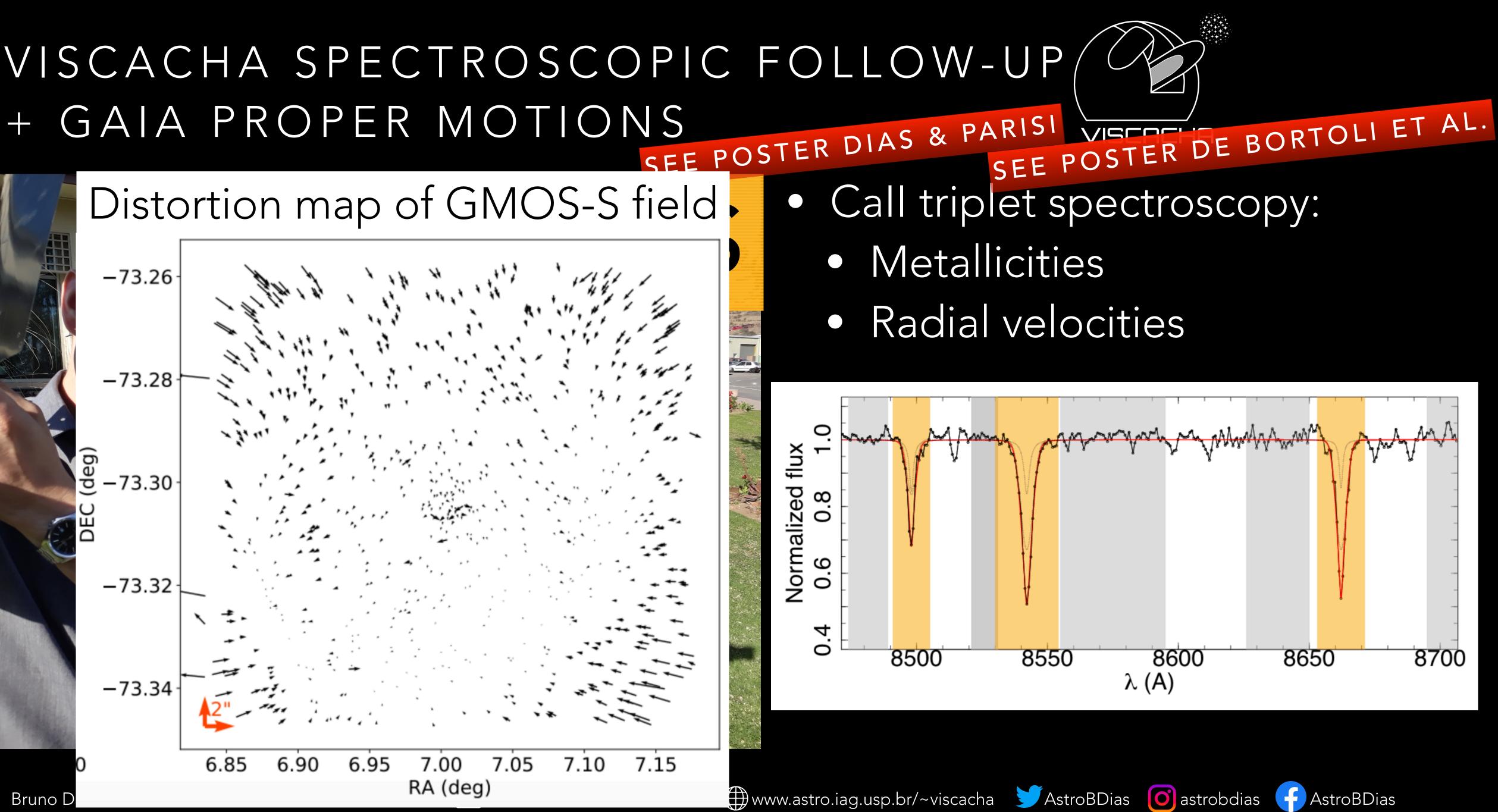






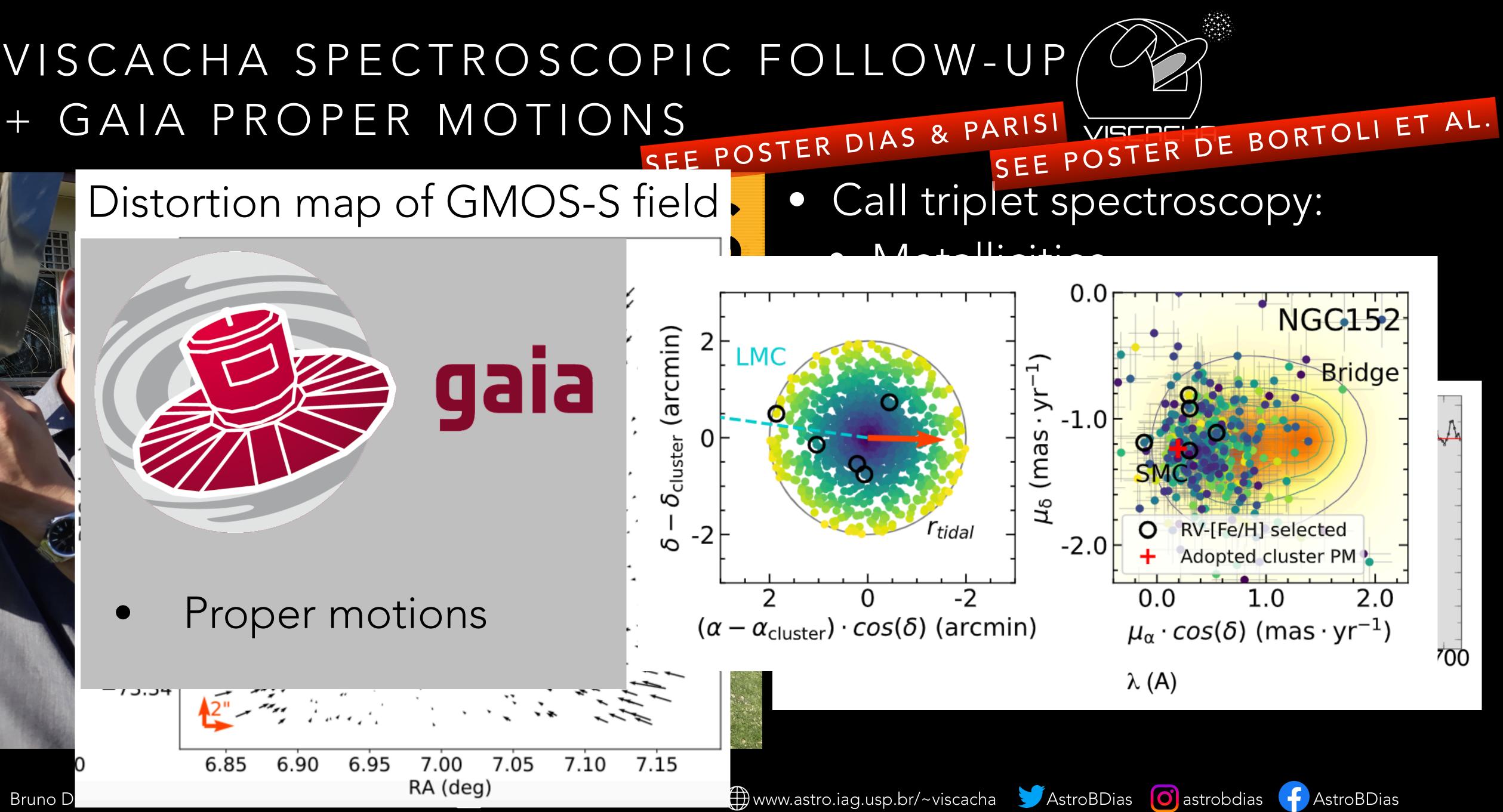


GAIA PROPER MOTIONS



Bruno D

+ GAIA PROPER MOTIONS



FURTHER EVIDENCE TO BE CONSTRAINED BY OBSERVATIONS

- Besla et al. (2012) analysed the case of a direct collision SMC-LMC
 - Model 1 has no collision and it does not affect much the SMC morphology

How do the star clusters see the SMC and LMC morphology?



Model 2 has collision and it produces the Bridge and Counter-Bridge from the SMC stars (and gas) and it also produces a warp in the LMC







SMC-LMC DIRECT COLLISION PRODUCES A WARP IN THE LMC

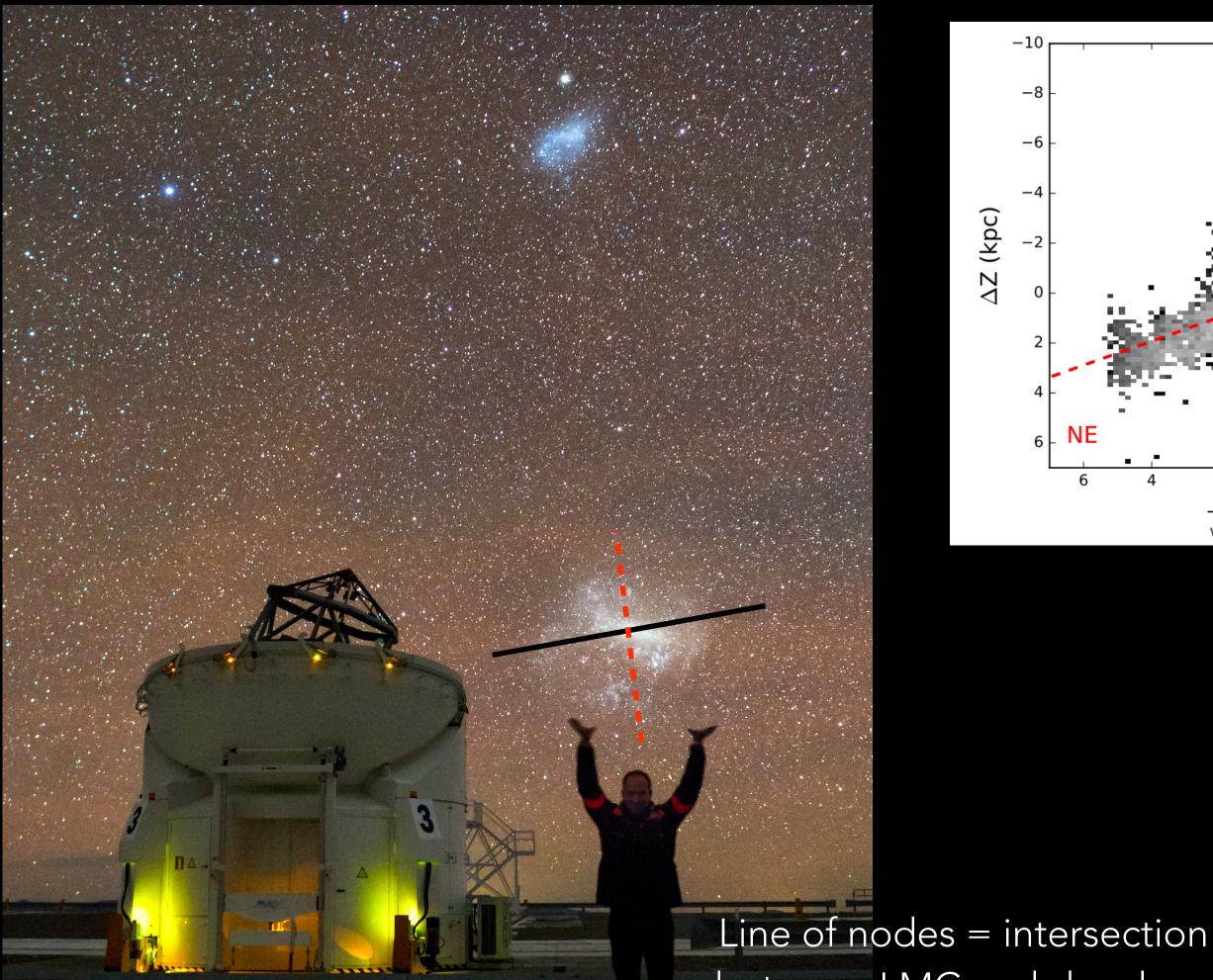
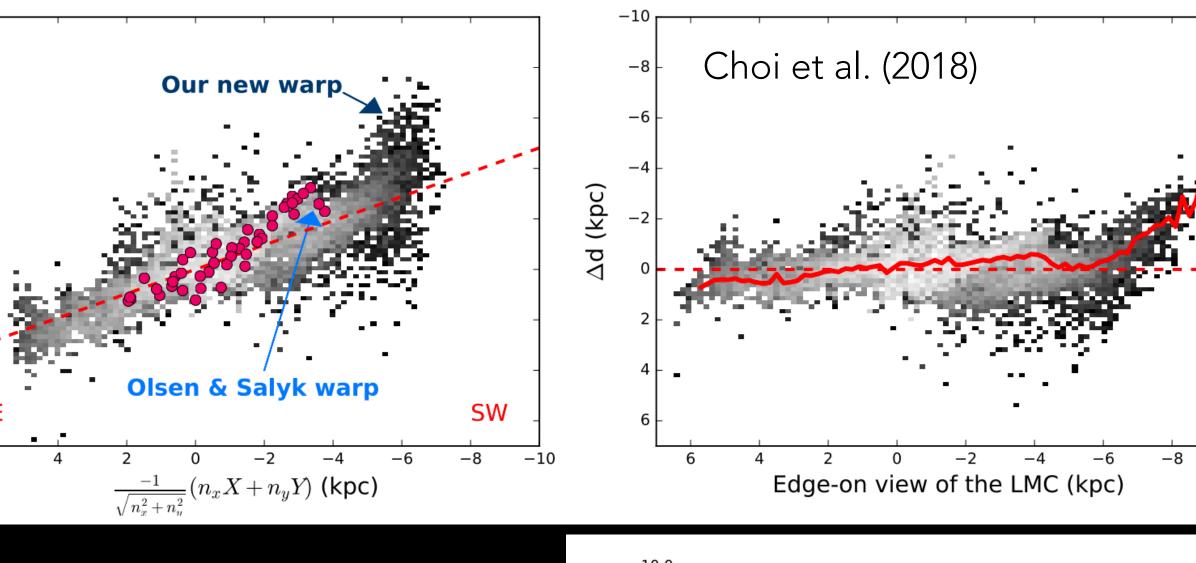


Photo by Juan Carlos Muñoz (ESO)

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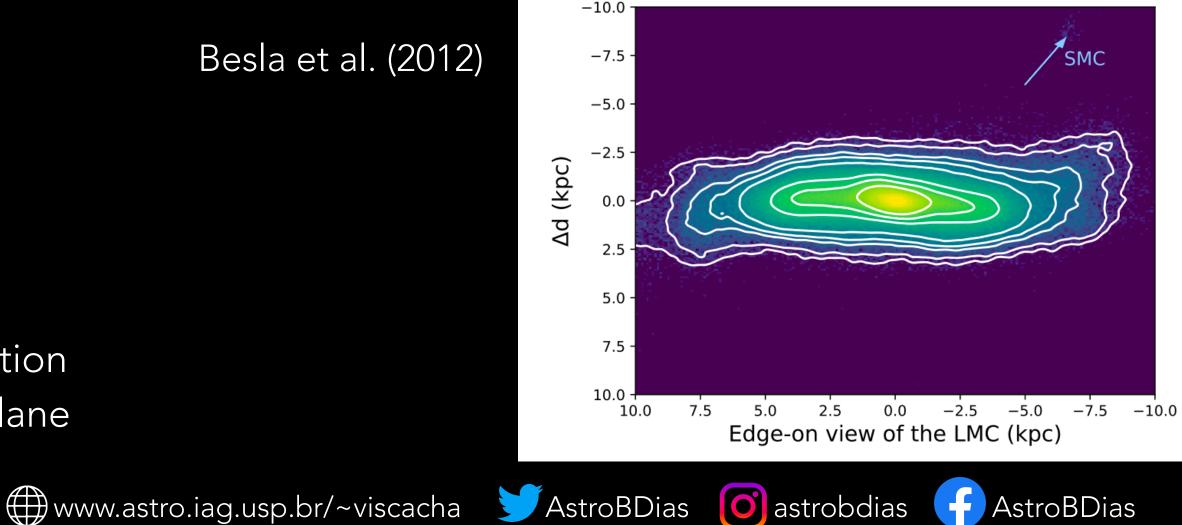






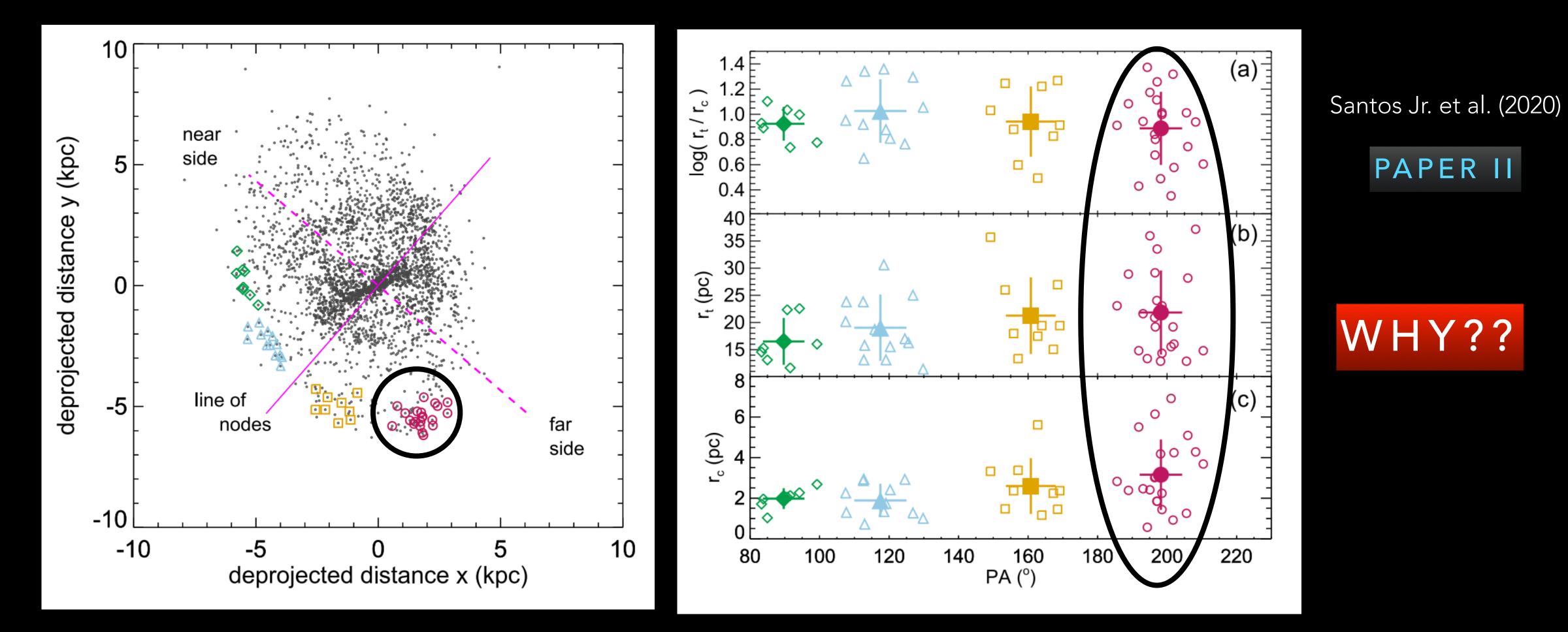
Besla et al. (2012)

between LMC and sky plane





STAR CLUSTERS IN THE WARP REGION SHOWS A LARGE SPREAD IN STRUCTURAL PARAMETERS: COINCIDENCE? NO REASONABLE EXPLANATION OR RELATION FOUND SO FAR...



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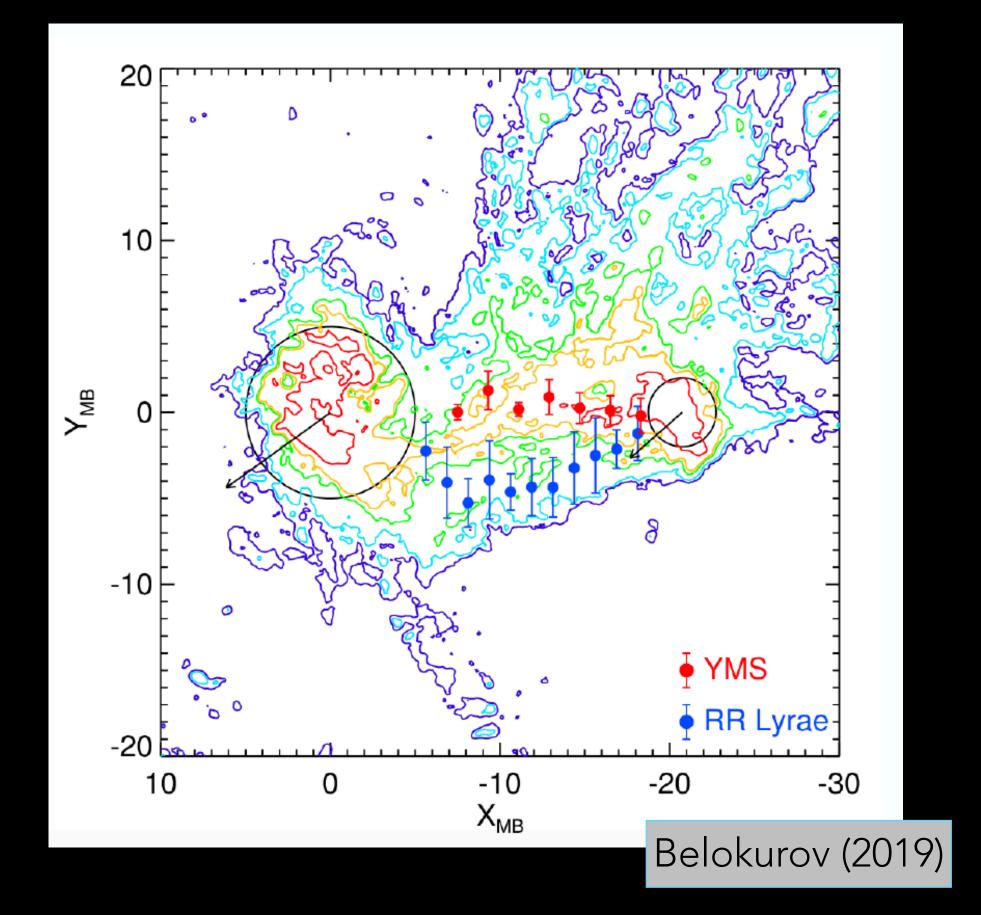








FIELD STARS



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STAR CLUSTERS



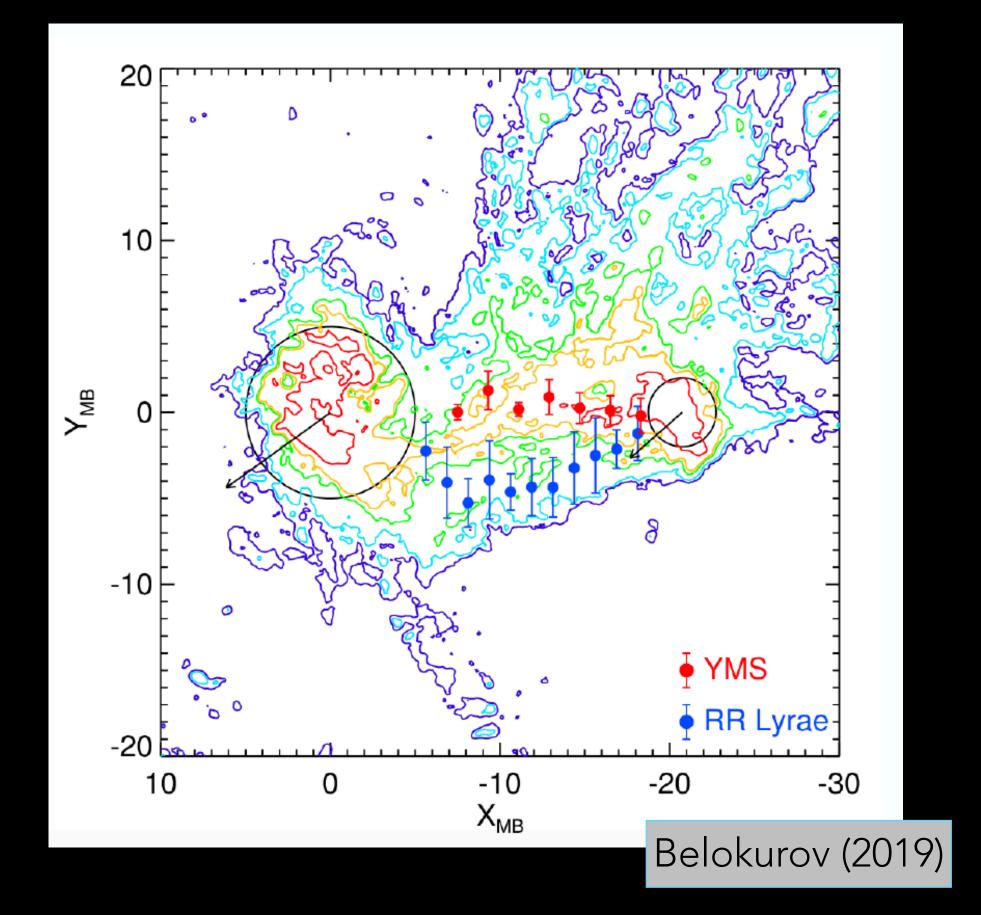








FIELD STARS

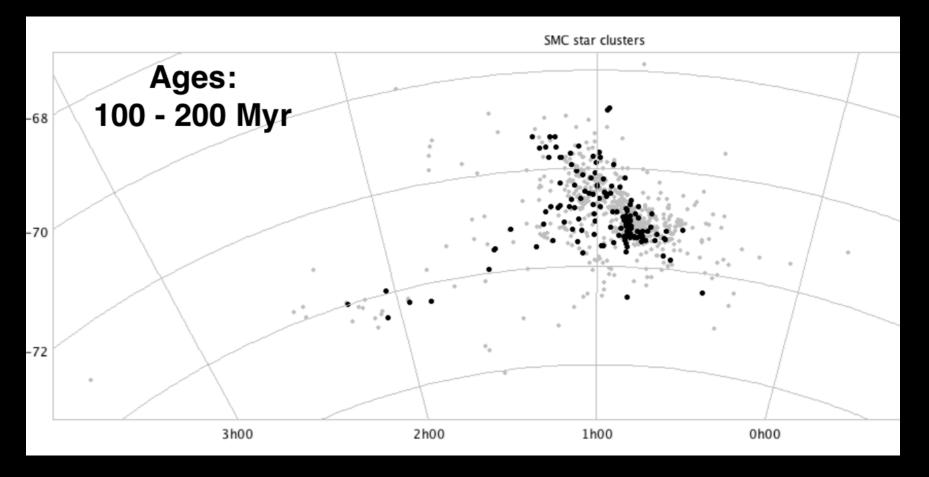


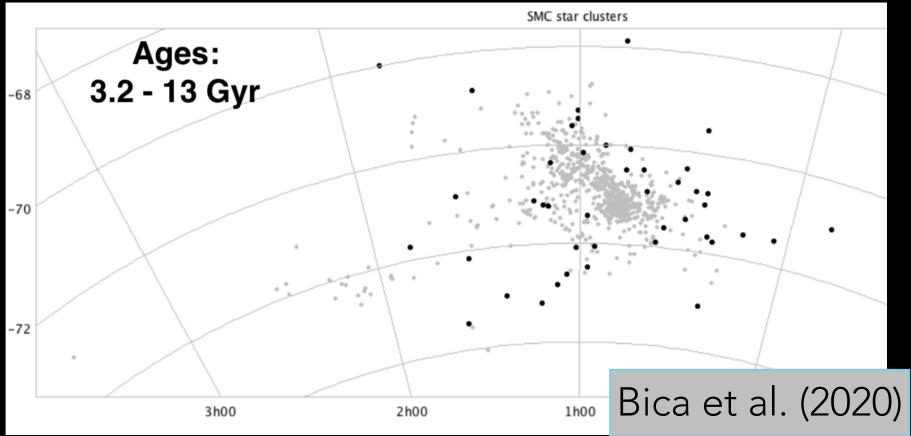
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STAR CLUSTERS



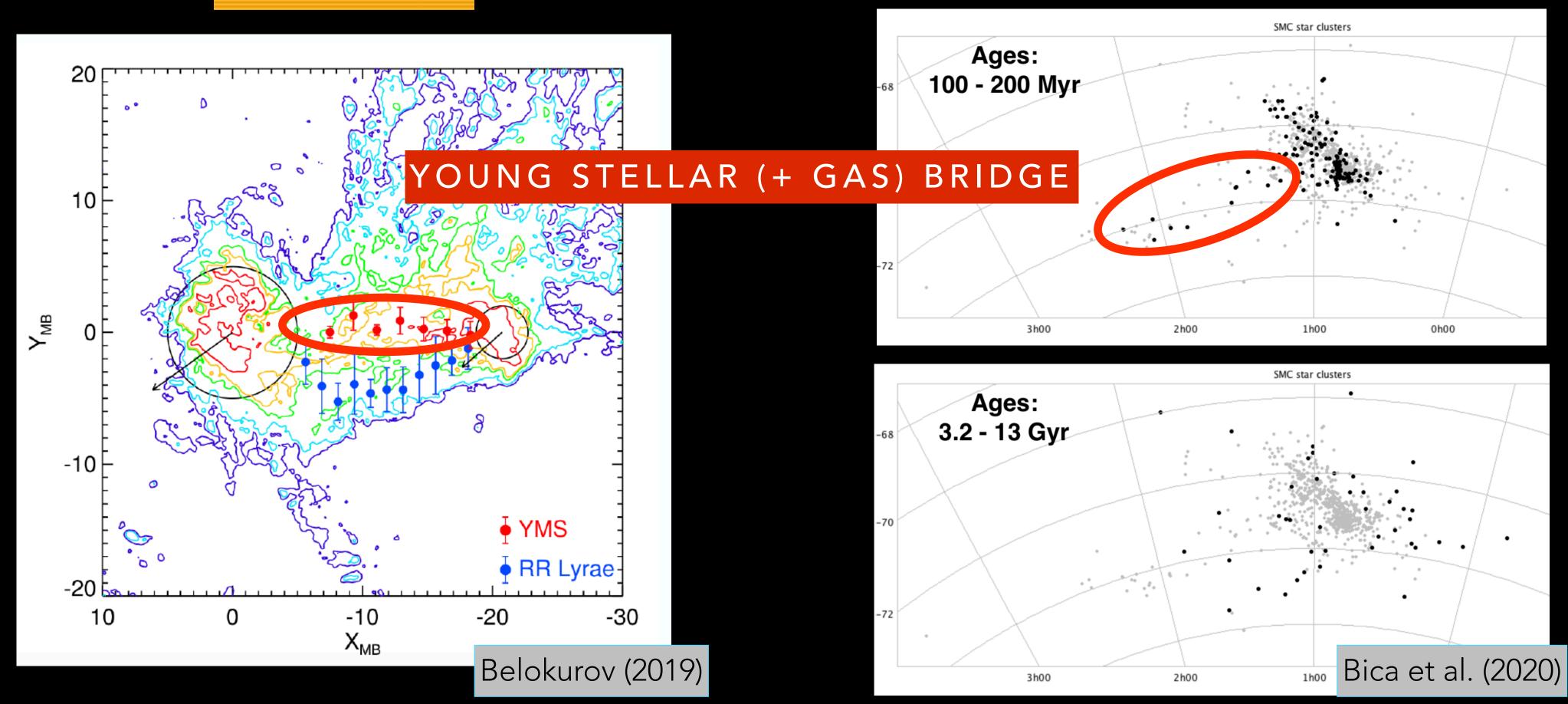








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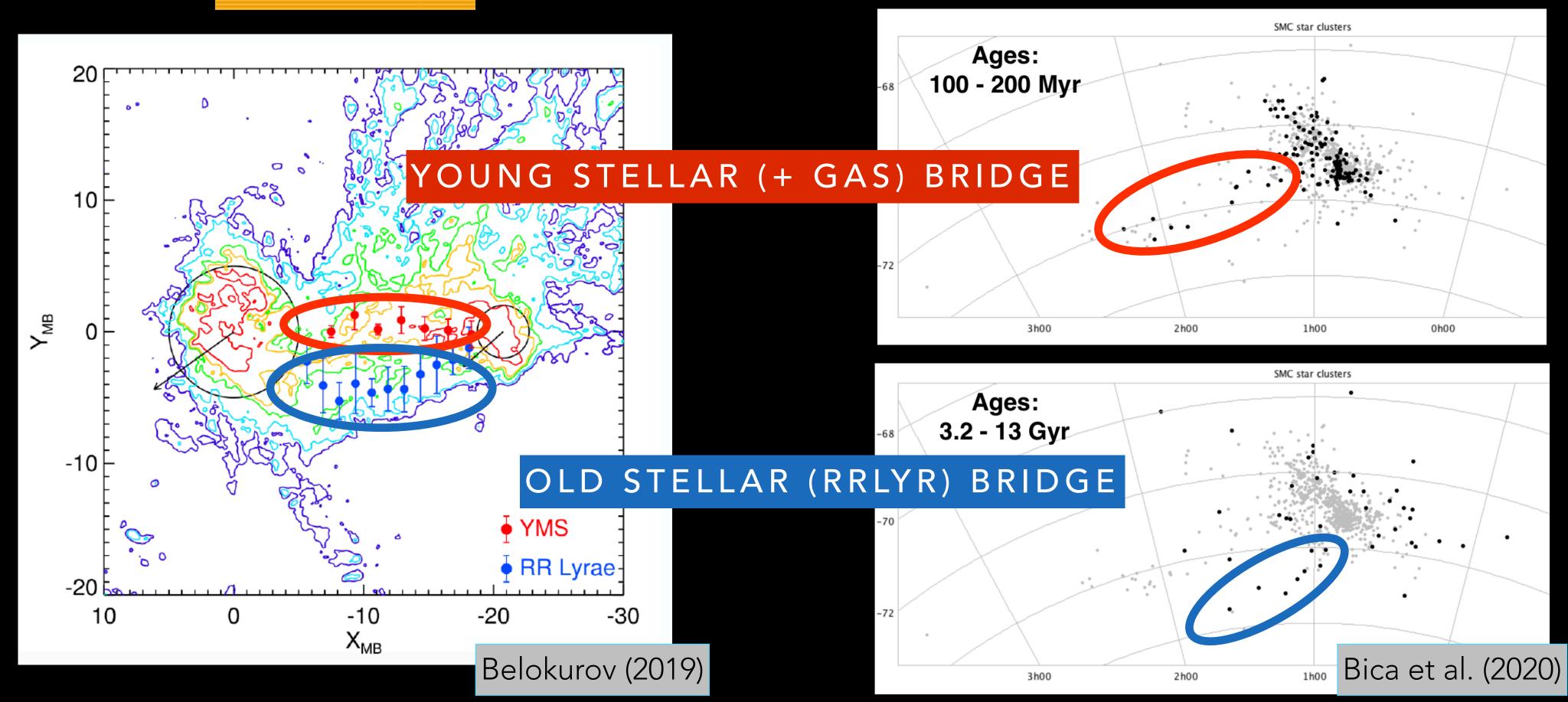
STAR CLUSTERS







FIELD STARS



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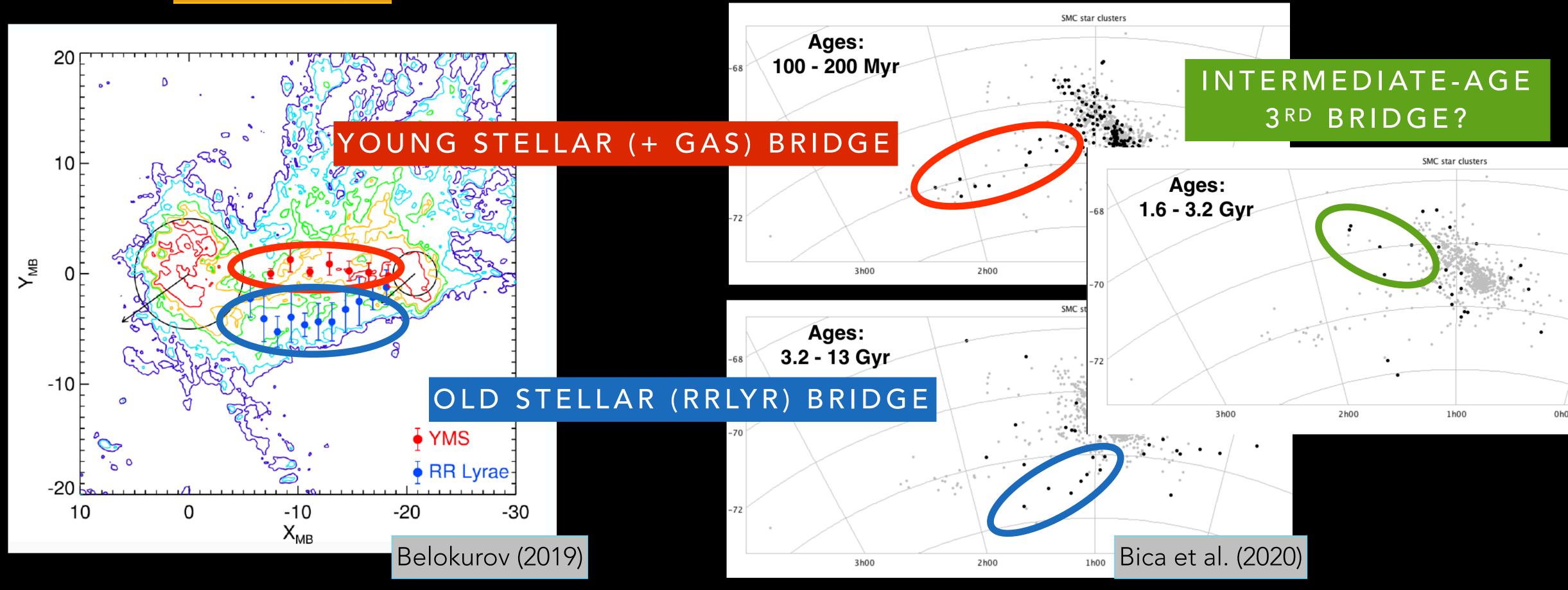


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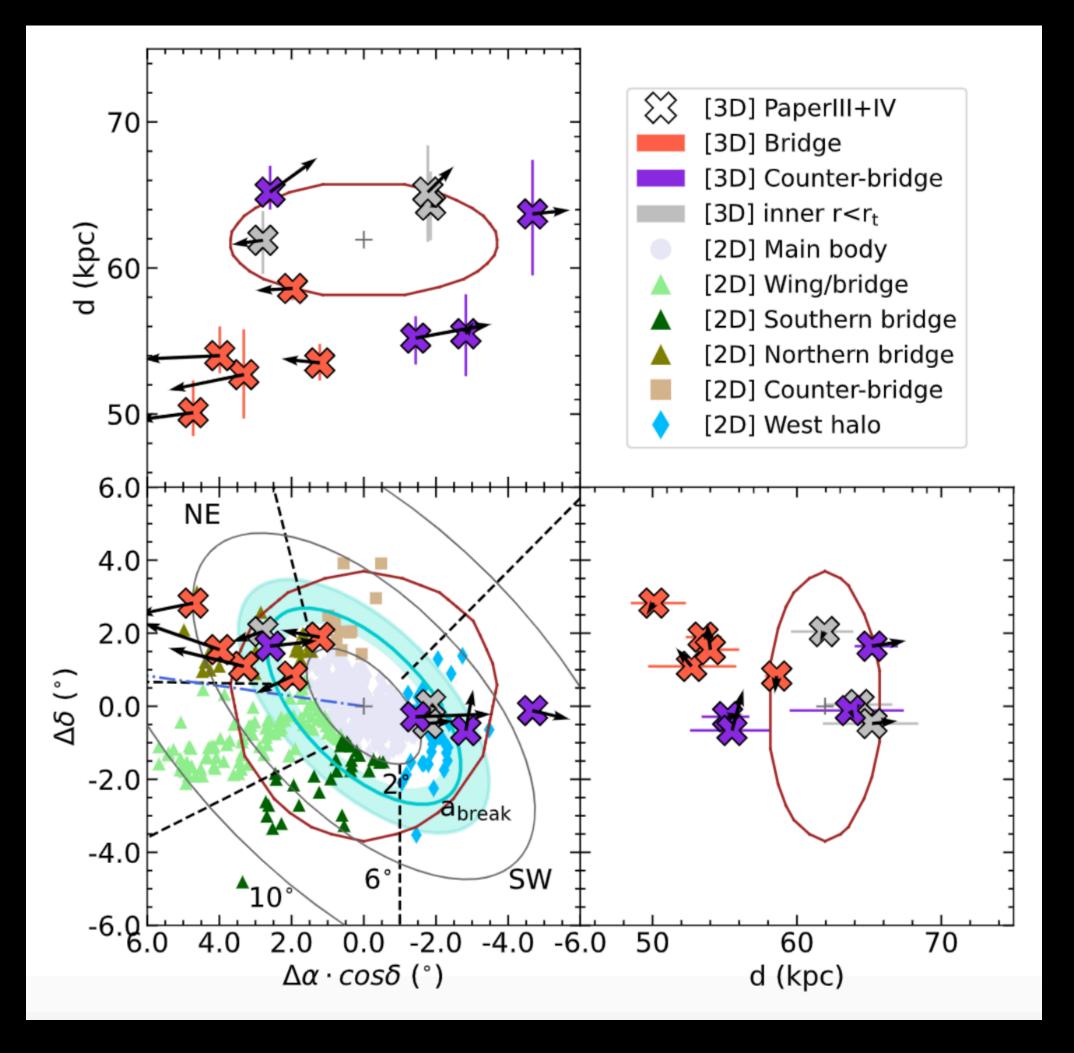
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A THIRD BRIDGE BRANCH DETECTED: IT IS ALSO OLD...

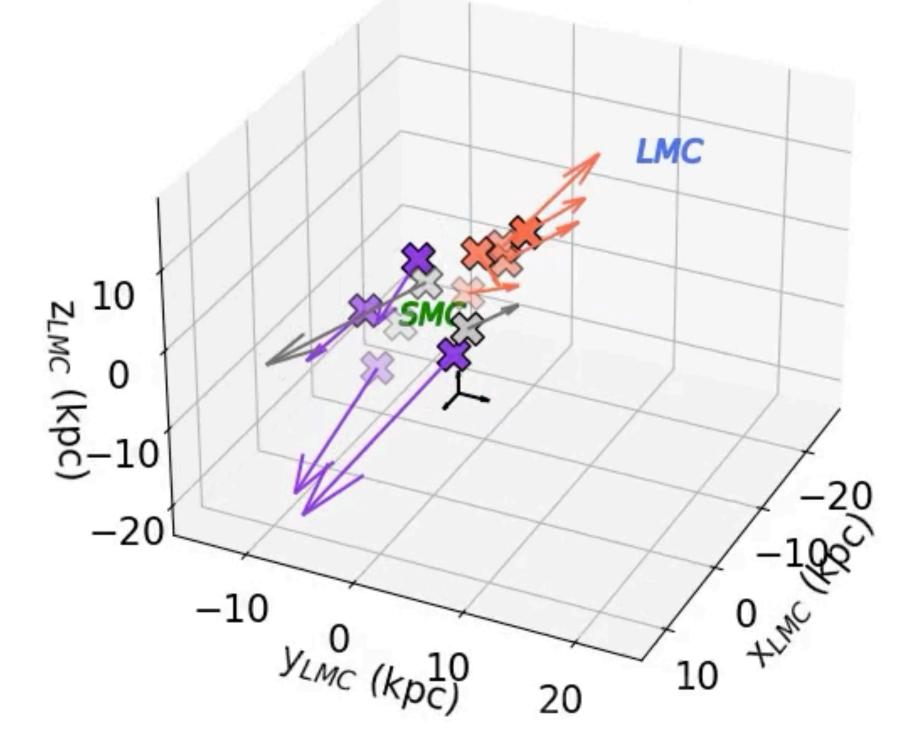


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GEMINIM was key to find these results

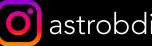




Dias et al. (2021,2022) **PAPER III, IV**









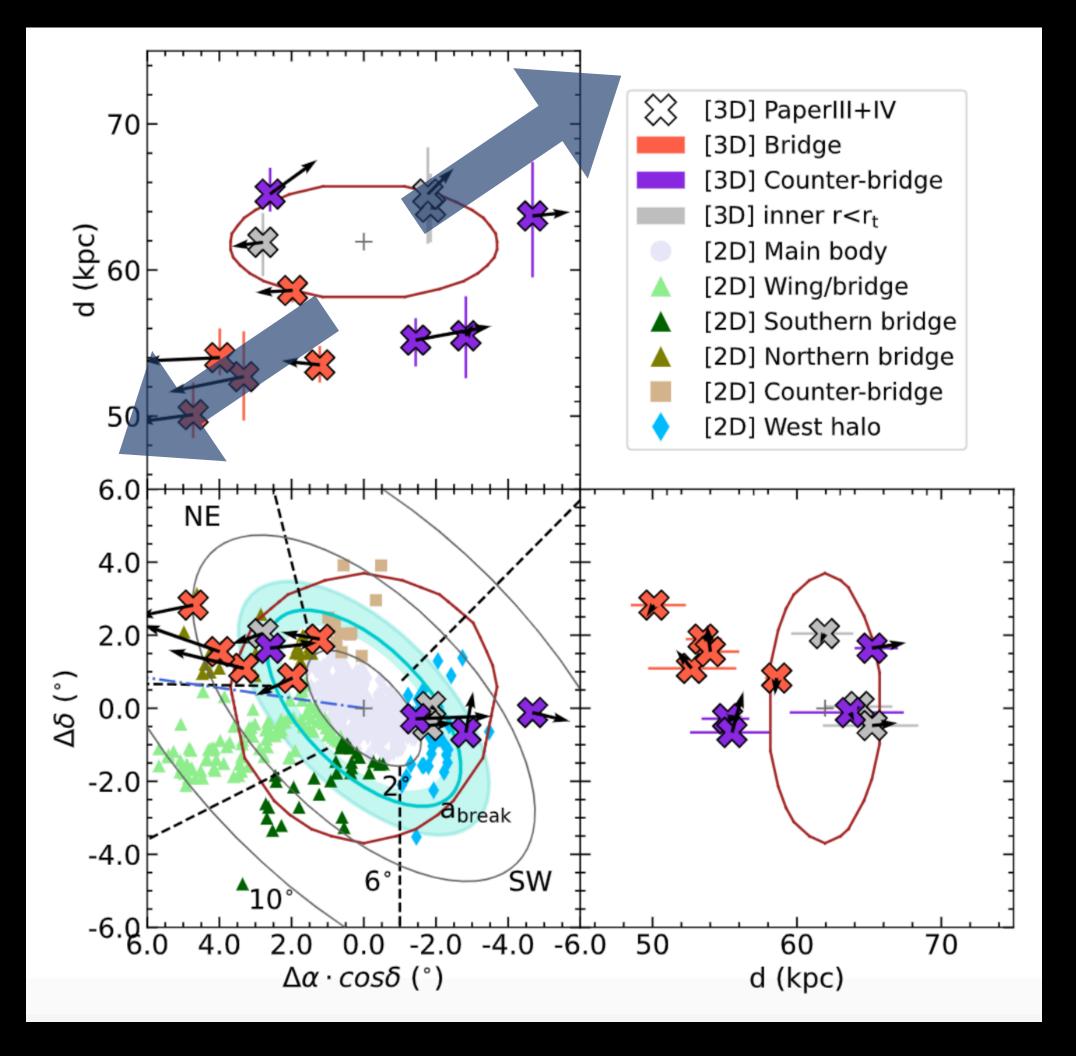








A THIRD BRIDGE BRANCH DETECTED: IT IS ALSO OLD...

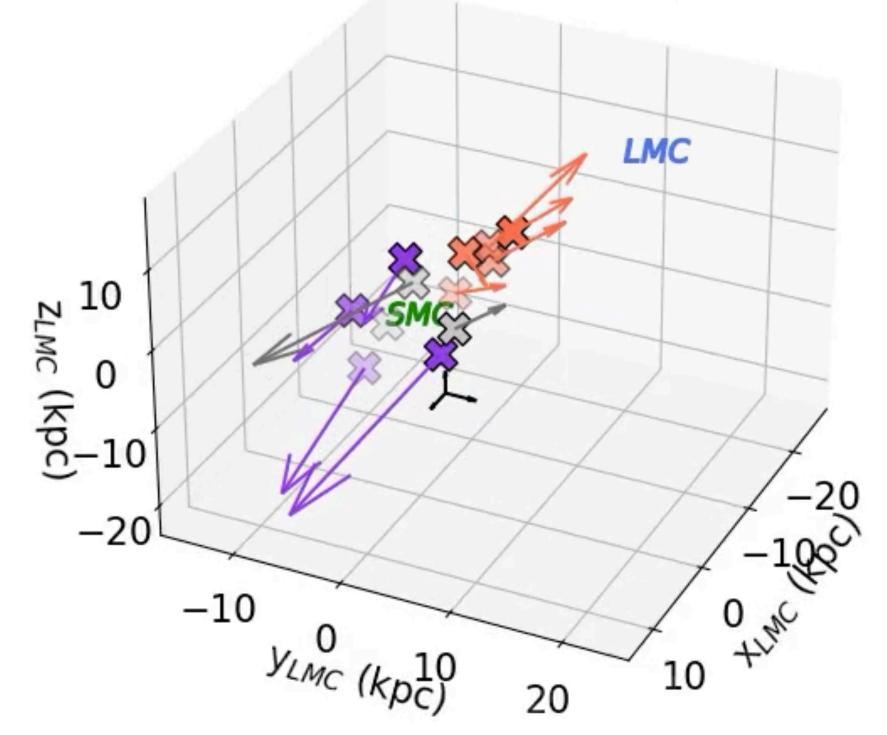


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GEMINE was key to find these results

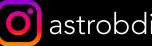




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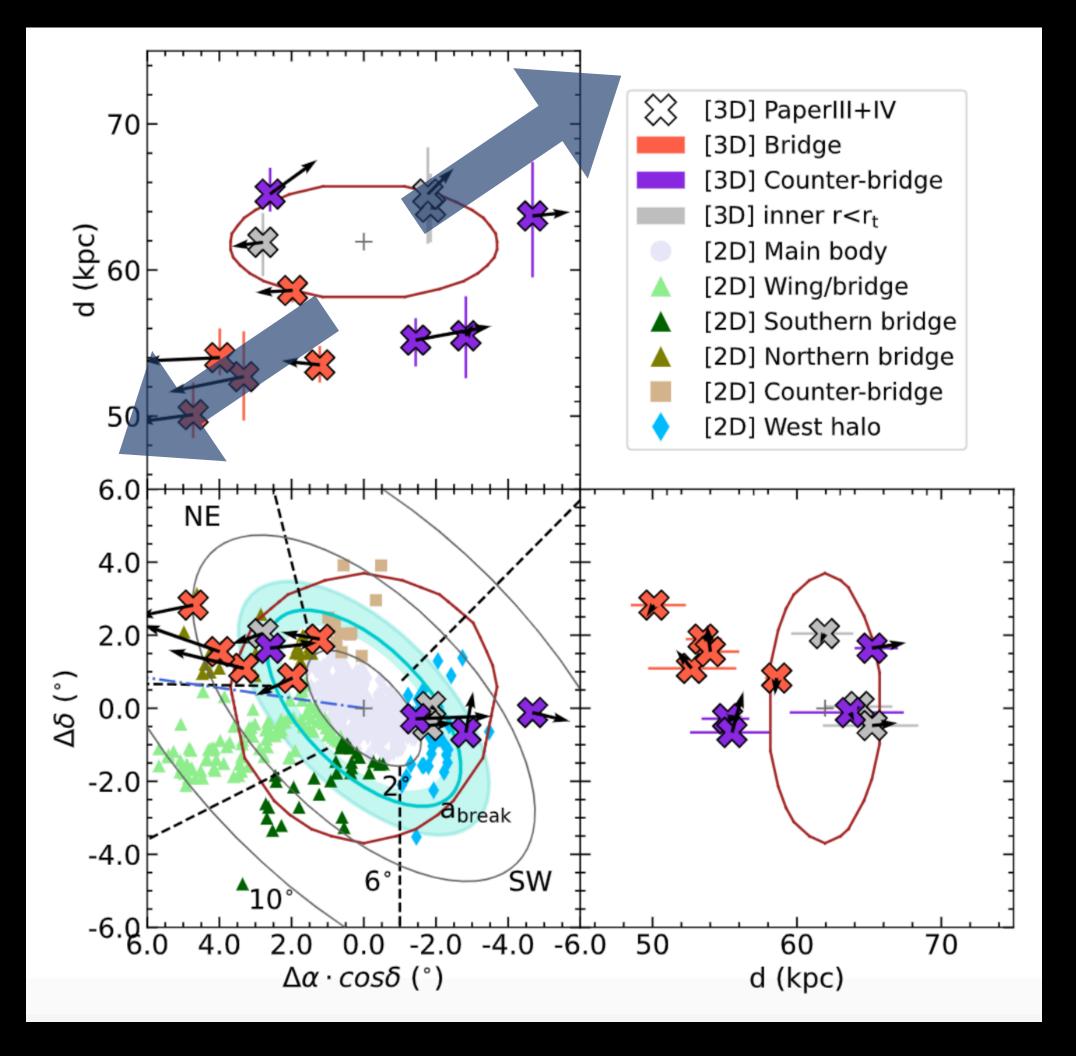








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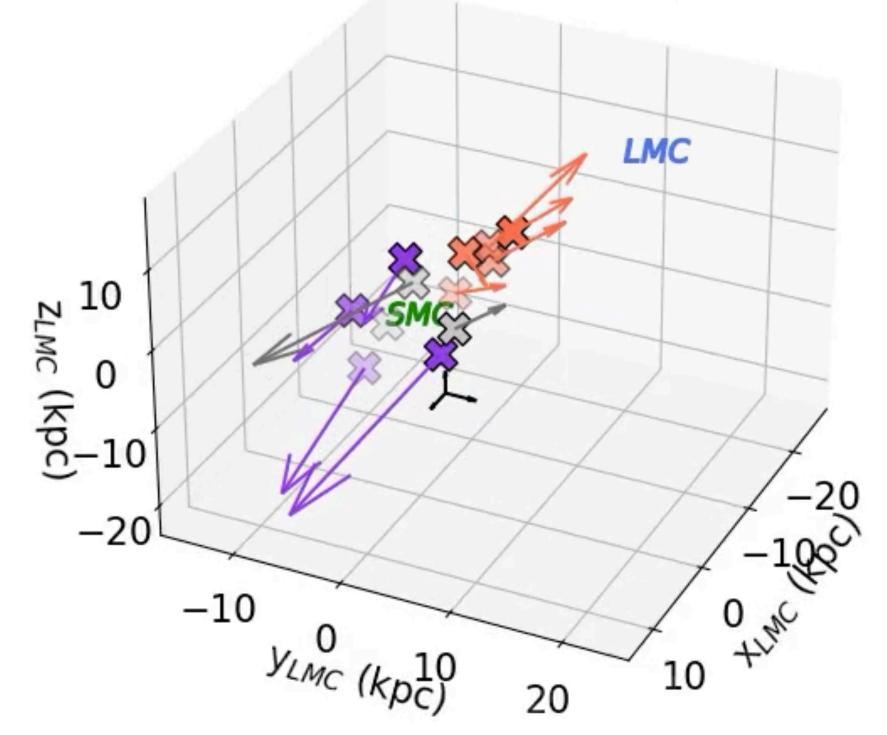


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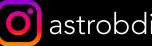




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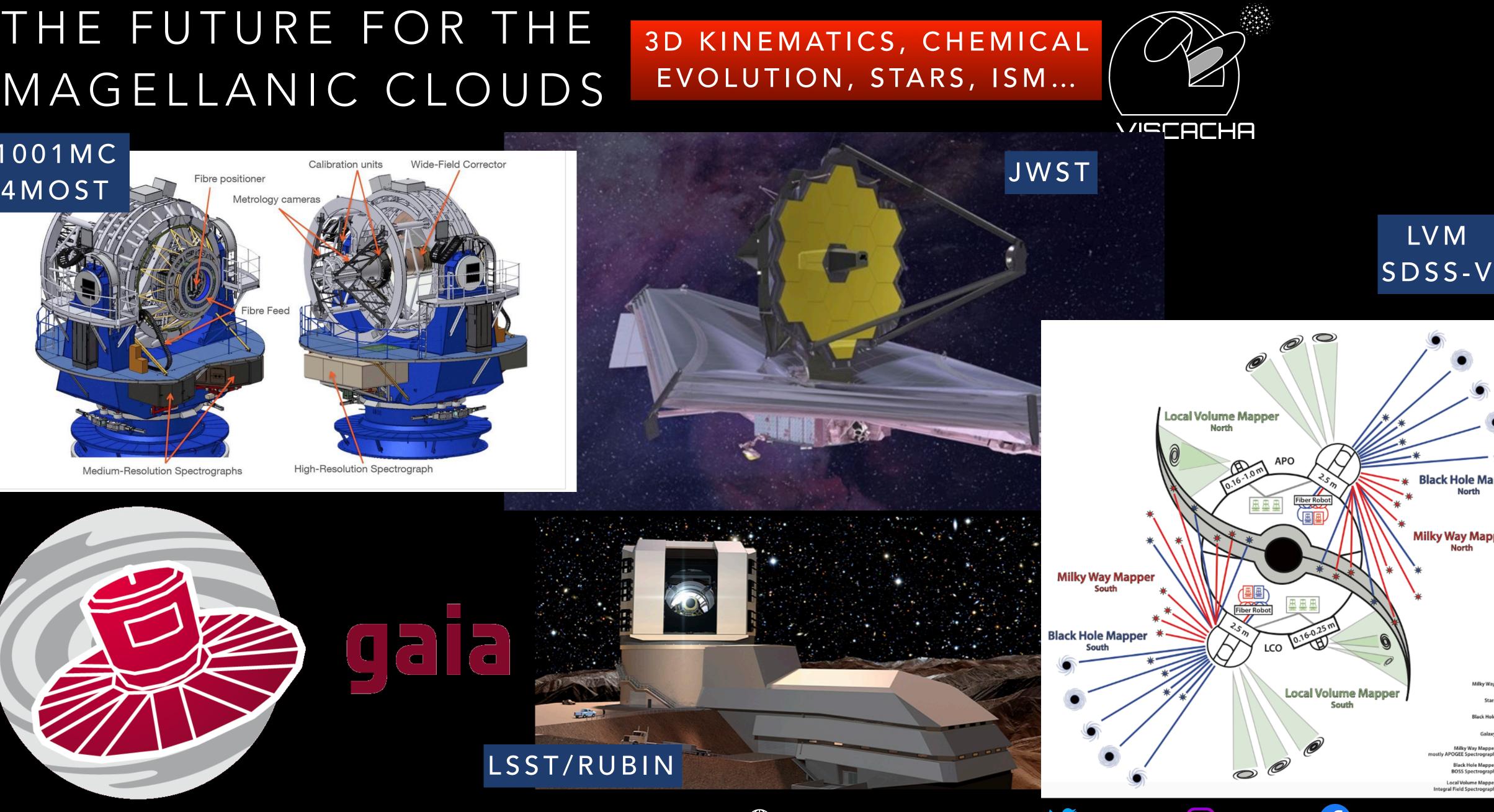


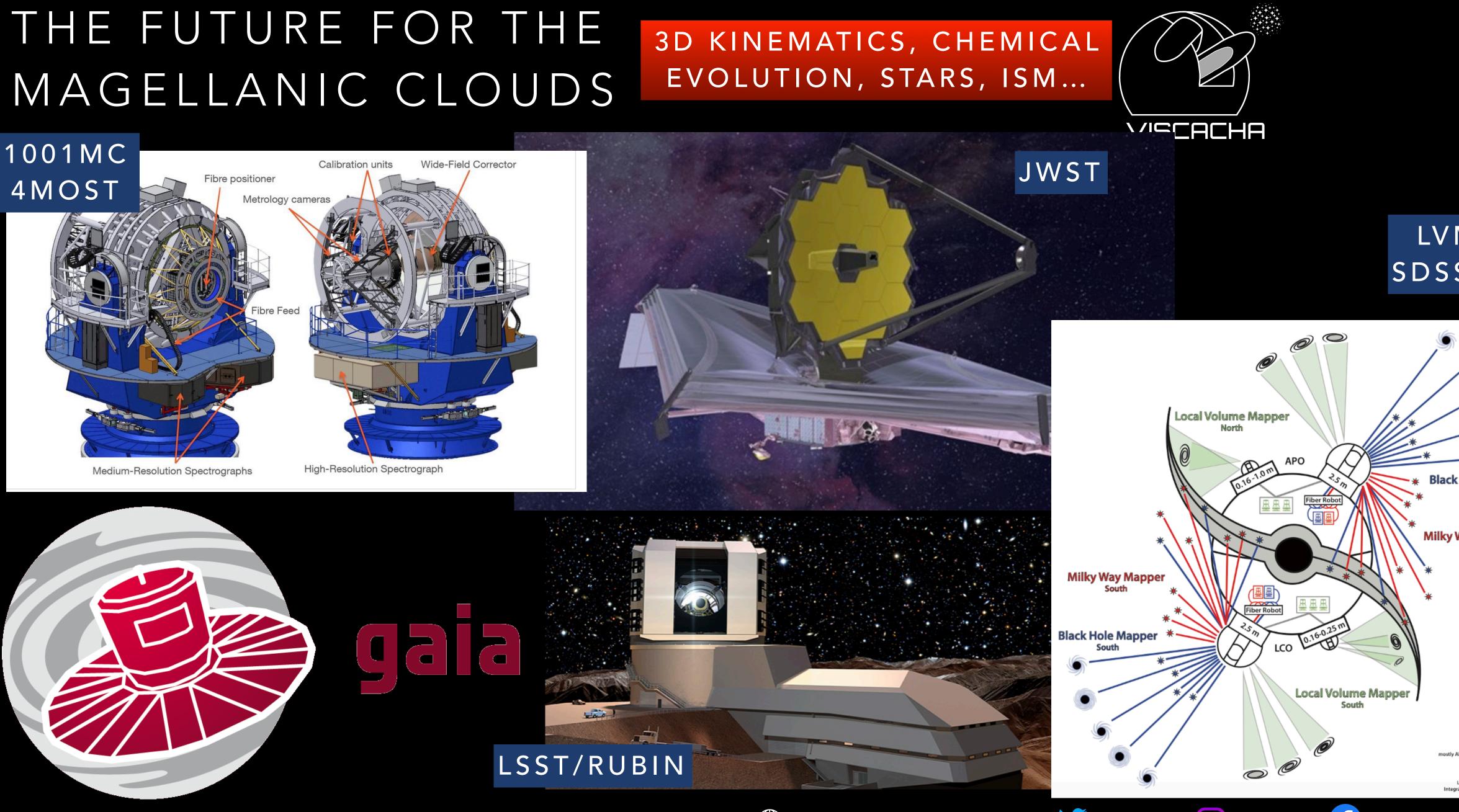








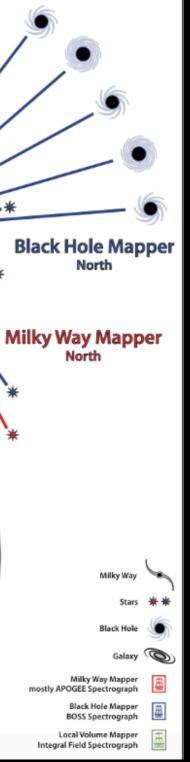


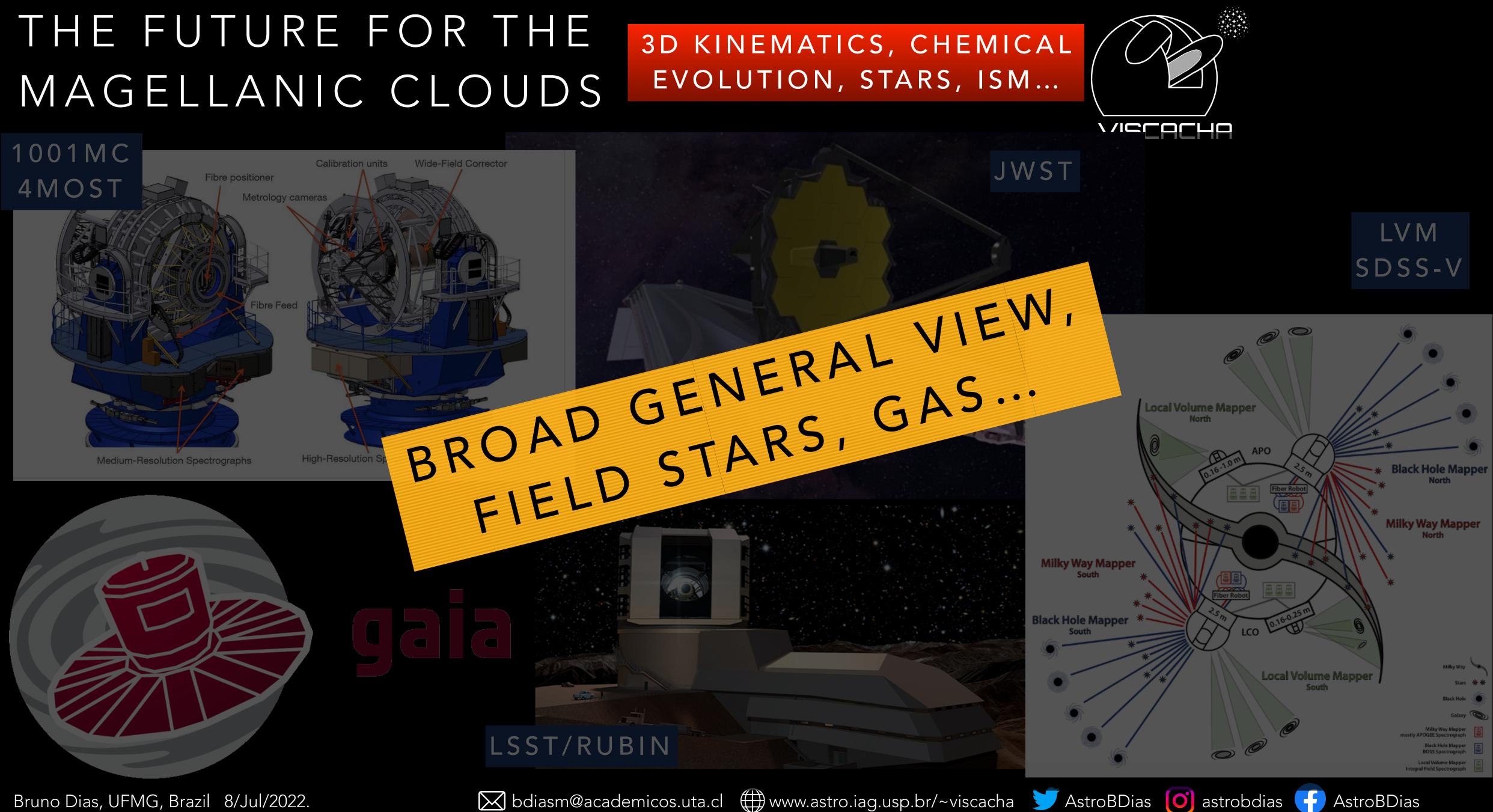


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THE FUTURE FOR THE MAGELLANIC CLOUDS



FOLLOW-UP DETAILED OBSERVATIONS FOR SPECIFIC TARGETS: STAR CLUSTERS

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- GMOS-S: continue our program on RV and [Fe/H] with CaT spectroscopy
- GMOS-S: variability at cluster core (see poster by Martinez-Vasquez)
- GSAOI+Gems: deeper CMDs for the star cluster cores









TAKE HOME MESSAGES

- The Magellanic Clouds have gained a lot of attention of large surveys and still too many discoveries to be done and to follow-up
- **VISCACHA** survey plays a key role wrt star clusters in the Magellanic Clouds
- The complex structure of the SMC and LMC are being enlightened by **VISCACHA** clusters
- Gemini is crucial to get cluster kinematics, metallicities, and even deeper CMDs in the cluster cores

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THANK YOU!

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Photo by Juan Carlos Muñoz (ESO)

